



1	Improved parameterization of snow albedo in WRF + Noah. Part II:
2	Applicability to snow estimates for the Tibetan Plateau
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19 Abstract

20	Snow albedo is important to the land surface energy balance and to the water cycle.							
21	During snowfall and subsequent snowmelt, snow albedo is usually parameterized as							
22	functions of snow related variables in land surface models. However, the default snow							
23	albedo scheme in the widely used Noah land surface model shows evident shortcomings							
24	in land-atmosphere interactions estimates during snow events on the Tibetan Plateau.							
25	Here, we demonstrate that our improved snow albedo scheme performs well after							
26	including snow depth as an additional factor. By coupling the WRF and Noah models,							
27	this study comprehensively evaluates the performance of the improved snow albedo							
28	scheme in simulating eight snow events on the Tibetan Plateau. The modeling results							
29	are compared with WRF run with the default Noah scheme and in situ observations.							
30	The improved snow albedo scheme significantly outperforms the default Noah scheme							
31	in relation to air temperature, albedo and sensible heat flux estimates, by alleviating							
32	cold bias estimates, albedo overestimates and sensible heat flux underestimates,							
33	respectively. This in turn contributes to more accurate reproductions of snow event							
34	evolution. The averaged RMSE relative reductions (and relative increase in correlation							
35	coefficients) for air temperature, albedo, sensible heat flux and snow depth reach 27 $\%$							
36	(5 %), 32 % (69 %), 13 % (17 %) and 21 % (108 %) respectively. These results							
37	demonstrate the strong potential of our improved snow albedo parameterization scheme							
38	for snow event simulations on the Tibetan Plateau. Our study provides a theoretical							
39	reference for researchers committed to further improving the snow albedo							
40	parameterization scheme.							

41 Keywords: WRF; snow albedo parameterization; turbulent heat and vapor fluxes;
42 Tibetan Plateau

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44 1 Introduction

The surface albedo directly determines the proportion of incident solar radiation that is absorbed by the surface, and is an important parameter in climate and land surface models (LSMs) (Sellers et al., 1996). Small changes in surface albedo can affect the energy balance in the land-atmosphere system, and can drive both local and global climate change (Bloch, 1964).

Surface albedo changes dramatically during snowfall and snowmelt cycles. Much 50 research has been carried out to identify the factors that influence these changes, 51 52 including the effects of terrain shielding, altitude, sky conditions, vegetation, and snow 53 properties such as grain size, liquid water content, depth, and impurities (Warren and 54 Wiscombe, 1980; Wiscombe and Warren, 1980; Aoki et al., 2003; Jonsell et al., 2003; 55 Hansen and Nazarenko, 2004; Liang et al., 2005; Wang et al., 2015; He et al., 2018a). 56 This body of research has led to the development of many parameterization schemes 57 for surface albedo (Oerlemans and Knap, 1998; Wang et al., 2007; Bao et al., 2008; Li and Hu, 2009; Gardner and Sharp, 2010; Kuipers Munneke et al., 2011; Malik et al., 58 59 2014; Dang et al., 2015; He et al., 2017, 2018b; Meng and Li, 2019; Saito et al., 2019; Wang et al., 2020). Most snow albedo parameterization schemes depend on statistical 60 empirical formulas and constant parameters, rather than representing physical snow-61 albedo feedback processes. To improve the performance of snow albedo 62 parameterization schemes for simulating land-atmosphere interactions, Bao and Lyu 63 (2009) added consideration of solar zenith angel to a regional climate model, which 64 resulted in a 1.2 °C temperature increase, and considerably improved the cold bias in 65 East Asia and improved the representation of diurnal ground temperature changes in 66 northwest China. Park and Park (2016) investigated the effect of vegetation on snow 67 68 covered surface albedo and improved the winter surface albedo estimates from their LSM by including leaf and stem indices in the snow albedo parameterization scheme, 69 which reduced the root mean square error (RMSE) by 69 %. Zhong et al. (2017) 70 considered aerosol radiative effects on snow processes in their simulations and 71





successfully reproduced the snow albedo and snow depth. Fresh snow albedo depends on snow depth, and albedo parameterization schemes that fail to account for this generally overestimate the snow depth. To address this, Wang et al. (2020) developed a new albedo scheme for fresh snow, which accounts for the relationship between fresh snow albedo, snow grain size and snow depth, resulting in improved snow depth estimates during the snow ablation period on the Tibetan Plateau. This highlights the importance of accounting for the effect of snow depth on fresh snow albedo.

A coupled land-atmosphere model can provide useful insights into conditions on the 79 Tibetan Plateau, where the terrain is complex and there are few, and unevenly 80 81 distributed observation stations (Maussion et al., 2011; Yuan et al., 2016; Norris et al., 2017; Bonekamp et al., 2018; Rahimi et al., 2019). However, the parameterization 82 scheme for surface albedo in the Noah LSM, which is currently the most widely used 83 LSM, does not account for all the factors that influence albedo. It includes many 84 predetermined parameters and an approximate treatment of vegetation, soil and snow, 85 which can result in some inaccuracies in the estimated surface albedo (Wen et al., 2011; 86 Liu et al., 2019). For example, the surface albedo parameterization scheme in the Noah 87 LSM considers snow cover and age, but ignores other snow related factors, such as 88 snow depth, that can drive dramatic changes in albedo (Ek et al., 2003). This makes it 89 inappropriate to use the Noah LSM to characterize changes in snow albedo that follow 90 from snowfall and melt processes in complex topographic areas. However, the Noah 91 LSM appears to be the most readily available snow albedo scheme for long term climate 92 modeling research (Rai et al., 2019). Despite its shortcomings, the Noah albedo 93 94 parameterization scheme does provide substantial improvements to estimates of the 95 magnitude and timing of both the peak snowfall amount and the maximum snow cover extent, following from the scheme's consideration of snow albedo decay and liquid 96 97 water refreezing (Livneh et al., 2010). The above issues represent opportunities for improvements to be made to the snow albedo parameterization scheme in the Noah 98 99 LSM.





100 The use of an advanced snow albedo parameterization scheme is crucial for accurate estimation of land-atmosphere interactions over the Tibetan Plateau, where the snow-101 albedo effect is extremely strong. It has been shown that the Weather Research and 102 Forecasting (WRF; Skamarock et al., 2008) model, when coupled with the default Noah 103 albedo parameterization scheme, results in an apparent cold bias over the Tibetan 104 Plateau (Gao et al., 2015; Meng et al., 2018; Liu et al., 2019). This bias can be reduced 105 including albedo products from the Moderate Resolution Imaging 106 bv Spectroradiometer (MODIS) and an additional snow depth parameter as independent 107 variables in the Noah albedo parameterization scheme (Liu, 2020). This approach is not 108 the same as assimilating satellite retrieved snow related products into the LSM, which 109 has also been shown to lead to improvements (Xu and Shu, 2014; Zhang et al., 2014; 110 Lin et al., 2016; Xue et al., 2019). This improved snow albedo scheme has been 111 successfully implemented in the WRF model, coupled with Noah, to simulate land-112 113 atmosphere interactions during a regional heavy snow event on the Tibetan Plateau (Liu, 114 2020). However, it has not been shown that the improvements that follow from the improved snow albedo scheme are universal over the Tibetan Plateau, and this should 115 116 be studied further. Severe snowfall occurs often over the southern Tibetan Plateau, while snowfall over the eastern Tibetan Plateau is generally of relatively weak intensity, 117 and the rate of snowmelt varies widely depending on the heterogeneous underlying 118 119 surfaces. This makes it necessary to carry out numerical experiments that focus on snow events over the eastern and southern Tibetan Plateau to assess how reliably the 120 121 improved scheme can be used to characterize different snowfall intensities and 122 snowmelt processes.

In this study, we selected eight moderate to snowstorm events that occurred over the southern and eastern Tibetan Plateau to assess the universality of the improvements offered by our improved snow albedo scheme in WRF coupled with the Noah LSM. For each snow event, two numerical experiments were carried out: one implementing the default Noah snow albedo scheme, and the other implementing our improved snow albedo scheme. The model performance was assessed through comparison of the





modeled air temperature, albedo, snow depth, turbulent heat and vapor fluxes with ground observations. The aim of this study is to explore the potential of our improved snow albedo parameterization scheme to simulate snow events over the whole Tibetan Plateau more accurately than can be done using the standard default scheme. We hope that this study will also provide a useful reference for researchers working to develop and improve this, and other albedo parameterization schemes.

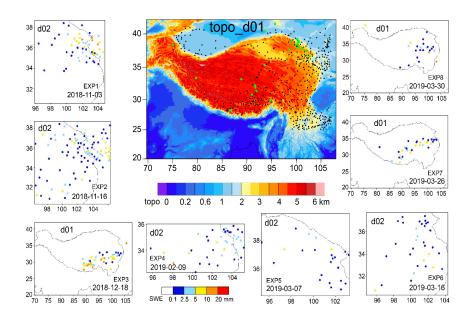
135 2 Data and methodology

136 2.1 Description of snow events

We selected China Meteorological Administration (CMA) national observation stations 137 on the Tibetan Plateau, and in the surrounding regions, with elevations exceeding 1000 138 m. This resulted in a total of 502 stations (Figure 1 shows their distribution). Snowfall 139 events were identified from the hourly air temperature and precipitation observations 140 141 from all 502 stations when the air temperature was below 0 °C. Daily snowfall amounts were calculated, using 08:00 Beijing Standard Time (BST) as the start and end time for 142 each day. The standards to define snowfall grade were taken from the China 143 Meteorological Standardization Network (http://www.cmastd.cn/). Using these 144 standards, eight different grades of snowfall event were considered in our study, 145 including moderate, heavy and snowstorm. Most of the snowfall events took place over 146 the eastern Tibetan Plateau, and some occurred in a large region across the southern and 147 central to the eastern Tibetan Plateau. The maximum daily snowfall amount from all 148 149 snowfall events exceeded 8 mm, and four events resulted in more than 10 mm daily 150 snowfall, making these snowstorm grade events. Snow depth is much greater on the 151 southern Tibetan Plateau (> 50 cm) than it is on the eastern and central Tibetan Plateau (<=20 cm). The description of eight snowfall events, including the date and location of 152 moderate to snowstorm grade events, the maximum snow depth and daily snowfall 153 amount, are detailed in Table 1. The daily snowfall amounts from the eight snowfall 154 events are shown in Figure 1. 155







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Figure 1. WRF domains of d01 and d02, and CMA observations of daily snowfall amount in color solid circles for the eight experiments. The topographical height of d01 is shaded with black solid circles indicating the locations of the CMA stations, and green solid circles indicating the locations of the CAS stations and the Qilian Mountains integrated observatory network.

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170 Table 1 Description of eight snowfall events using China Meteorological

171 Administration observations.

Snow events	Event 1	Event 2	Event 3	Event 4	Event 5	Event 6	Event 7	Event 8
Date of moderate to snowstorm	Nov. 3-6, 2018	Nov. 15-17, 2018	Dec. 17-18, 2018	Feb. 9, 2019	Mar. 7-8, 2019	Mar. 16, 2019	Mar. 26, 2019	Mar. 30-Apr. 1, 2019
Region in the Tibetan Plateau	eastern	eastern	southern, central, eastern	eastern	eastern	eastern	southern, central, eastern	central, eastern
Maximum daily snowfall amount (unit: mm)	13.6	9.2	18.7	11	9.3	9.1	8.7	14.3
Maximum snow depth (unit: cm)	18	16	65	20	13	9	53	18

172 2.2 Model description and experiment configuration

The WRF model (Skamarock et al., 2008), version 3.7.1, coupled with the Noah LSM, was used to simulate the eight snowfall events in this study. It is a fully compressible, non-hydrostatic model and includes a run time hydrostatic option. Vertical levels are determined using a mass based terrain following hydrostatic pressure coordinate, and calculations are performed on an Arakawa C grid. The model uses 2nd and 3rd order Runge-Kutta time integration schemes, and 2nd to 6th order advection schemes in both the horizontal and vertical direction.

The extremely steep terrain on the central and southern Tibetan Plateau led to model instability and failure for snowfall events 3, 7 and 8 when a relatively fine horizontal resolution of 1 km was used; however, the calculations remained stable when the resolution was increased to 5 km. We therefore used two two ways nested modeling





184	domains in our model configuration for snowfall events 1, 2, 4, 5 and 6, and a single
185	modeling domain for snowfall events 3, 7 and 8. The coarse domain (d01) was used to
186	simulate synoptic scale atmospheric conditions over 20.0-42.0° N and 69.7-108.0° E
187	with a horizontal resolution of 5 km. The inner domain (d02) had a horizontal resolution
188	of 1 km, and event 1 occupied 876×966 grid cells, event 2 occupied 976×1001 cells,
189	event 4 was resolved by 966 \times 451 cells, event 5 was calculated over 781 \times 686 cells,
190	and event 6 covered 926 \times 881 cells. The vertical structure of both domains included
191	35 unevenly spaced layers and extended up to 50 hPa. The model was configured to use
192	the Noah LSM in d01 and d02 to describe all land-atmosphere interactions; the
193	Thompson scheme to represent microphysical processes; the Dudhia scheme to
194	represent shortwave radiation, the RRTM scheme to describe longwave radiation; the
195	YSU scheme to describe the planetary boundary layer; and only in d01 to use the Kain-
196	Fritsch cumulus parameterization scheme for clouds.

We conducted numerical experiments to simulate snow event 1 (EXP1), event 2 (EXP2), 197 event 3 (EXP3), event 4 (EXP4), event 5 (EXP5), event 6 (EXP6), event 7 (EXP7) and 198 199 event 8 (EXP8). The model domains for the experiments are shown in Figure 1. Each 200 experiment included two model simulations: one implementing the default Noah snow albedo parameterization scheme, and the other implementing a new improved snow 201 albedo scheme. The snow albedo is parameterized using Eq. (1) and (2) in the default 202 Noah land surface scheme (Livneh et al., 2010), and using Eq. (3) and (4) in the 203 improved scheme (Liu, 2020): 204

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$$\begin{aligned} \alpha_{\rm snow} &= \alpha_{max} \times A^{t^{\rm D}} \\ \alpha &= \alpha_{bg} + sc \times (\alpha_{\rm snow} - \alpha_{bg}) \end{aligned} \tag{1}$$

$$\alpha_{\rm snow} = 0.13 + 0.66e^{\left(\frac{t}{1.38}\right)}$$
(3)

$$\alpha = \alpha_{\rm snow} + (0.19 - \alpha_{\rm snow})e^{\left(\frac{-d}{0.11}\right)}$$
(4)

208 where A and B are constants, equal to 0.94 and 0.58, respectively, for snow





accumulation periods, and are 0.82 and 0.46, respectively, for other periods; α_{bg} is the background albedo, which depends on the land cover type; *sc* is snow cover, and ranges from 0 to 1; α_{max} is fresh snow albedo; α_{snow} is snow albedo; *t* is the snow age in units of days; *d* is snow depth in meters.

The fifth generation European Centre for Medium Range Weather Forecasts (ECMWF) 213 214 reanalysis dataset (ERA5), with a spatial resolution of $0.25^{\circ} \times 0.25^{\circ}$ and 3 h temporal resolution, provided initial and boundary conditions for our numerical experiments. The 215 ERA data were calculated using 4DVar data assimilation in CY41R2 from ECMWF's 216 Integrated Forecast System, with 137 vertical hybrid sigma/pressure levels, extending 217 218 0.01 hPa. ERA5 is freely available from the website to https://www.ecmwf.int/en/forecasts/datasets/reanalysis-datasets/era5. The near real 219 220 time MODIS land cover product was used to replace the outdated land cover in WRF preprocessing system. We ran the model from before the onset of each snowfall event 221 until and after all snowmelt following the event had ceased. EXP1 was run Nov. 1-8 222 223 2018, EXP2 was run Nov. 13-18 2018, EXP3 was run Dec. 16-20 2018, EXP4 was run Feb. 5-11 2019, EXP5 was run Mar. 2-10 2019, EXP6 was run Mar. 12-18 2019, EXP7 224 was run Mar. 24-29 2019, EXP8 was run Mar. 28-Apr.7 2019. The model results were 225 output at 3 hours intervals and the first day was used for model spin up. 226

227 2.3 Data for model evaluation and comparison

228 CMA hourly observations of air temperature and snow depth from 502 stations were 229 used to assess the WRF model estimates of 2 m air temperature and snow depth that 230 were made using the improved snow albedo parameterization scheme. Albedo is a key 231 factor for net radiation calculations, and is defined as the ratio of reflected shortwave radiation (upwards) to received shortwave radiation (downwards). It determines the 232 distribution of turbulent land surface heat fluxes between sensible (SH) and latent heat 233 234 (LH). There are many meteorological observations available that have been continuously recorded on the Tibetan Plateau at atmospheric boundary layer towers, 235 236 eddy covariance systems (Ma et al., 2018, 2020) and the Qilian Mountains integrated





237	observatory network (Li, 2019; Liu et al., 2020; Zhao and Zhang, 2020). These provide
238	in situ data that are assumed to constitute 'truth' for the model validation in this study.
239	In situ observations of albedo, SH and LH from 11 Chinese Academy of Sciences (CAS)
240	stations/samples, and from 16 stations in the Qilian Mountains integrated observatory
241	network are used to evaluate the accuracy of the improved snow albedo
242	parameterization scheme for modeling snow events on the Tibetan Plateau. It is
243	reasonable to compare observations of albedo, SH and LH with model estimates at 5
244	km resolution because there are only a few in situ observation stations in d02, but a
245	total of 27 observation stations in d01. At local solar noon in Lhasa (14:00 BST), the
246	observed albedo value is closer to the Lambertian albedo that is described by the WRF
247	model when coupled with LSMs. We therefore used albedo observations made at 14:00
248	BST to evaluate the model calculated albedo. Quality control codes 1, 2 and 3 were
249	selected when using the Turbulence Knight version 3 (TK3) software, and 0 was used
250	when using the Eddypro software to calculate SH and LH. Details of the 27 stations
251	from the CAS and the Qilian Mountains integrated observatory network that were used
252	in our study are provided in Table 2, and their locations on the Tibetan Plateau are
253	shown in Figure 1. The RMSE and the correlation coefficient were calculated for the
254	comparisons between the observations and the WRF estimates for albedo, air
255	temperature, SH, LH, and snow depth. These are used to assess the performance of the
256	WRF model, when implemented with the default or the improved snow albedo scheme.





- 263 Table 2 Location of stations from the Chinese Academy of Sciences (CAS) and the
- 264 Qilian Mountains integrated observatory network, and whether or not observations of
- albedo, SH and LH were used from each station.

Station No.	Station Name	Station Type	Latitude (° N)	Longitude (° E)	Elevation (m)	Using albedo	Using SH and LH
1	SETS	CAS	29.77	94.73	3326	Yes	No
2	QOMS	CAS	28.21	86.56	4276	Yes	Yes
3	QOMS sample	CAS	28.31	86.85	4600	No	Yes
4	MASWE	CAS	38.41	75.05	3668	Yes	No
5	Nam Co	CAS	30.77	90.99	4730	Yes	Yes
6	NASDE	CAS	33.39	79.70	4264	Yes	No
7	Shuanghu	CAS	33.22	88.83	4947	Yes	Yes
8	NewD66	CAS	35.43	93.59	4465	Yes	No
9	BJ	CAS	31.37	91.90	4509	Yes	Yes
10	MS3478	CAS	31.93	91.71	4620	Yes	No
11	Amdo	CAS	32.24	91.62	4695	Yes	No
12	Yakou	Qilian	38.01	100.24	4148	Yes	Yes
13	Arou	Qilian	38.05	100.46	3033	Yes	Yes
14	Jingyangling	Qilian	37.84	101.12	3750	Yes	Yes
15	Dashalong	Qilian	38.84	98.94	3739	Yes	Yes
16	Heihe Remote Sensing	Qilian	38.83	100.48	1560	Yes	No
17	Huazhaizi Desert Steppe	Qilian	38.77	100.32	1731	Yes	Yes
18	Daman	Qilian	38.86	100.37	1556	Yes	Yes
19	Zhangye wetland	Qilian	38.96	100.45	1460	Yes	Yes
20	Guazhou	Qilian	41.41	95.67	2014	Yes	Yes
21	Dayekou	Qilian	38.56	100.29	2703	Yes	No
22	Dunhuang	Qilian	40.35	93.71	993	Yes	No
23	Liancheng	Qilian	36.69	102.74	2903	Yes	No
24	Linze	Qilian	39.24	100.06	1402	Yes	No
25	Sidalong	Qilian	38.43	99.93	3146	Yes	No
26	Xiyinghe	Qilian	37.56	101.86	3616	Yes	No
27	Tianjun	Qilian	37.70	98.61	3718	Yes	Yes





266 3 Results

267 **3.1 Air temperature**

268 Albedo is a key factor in determining the net radiation received at the surface, which determines the land surface energy balance and influences air temperature. Scatterplots 269 of the air temperatures estimated by the WRF model and observed at the CMA stations 270 are shown in Figure 2. In all eight modeling experiments, implementing the improved 271 snow albedo scheme in the WRF model greatly reduces the cold bias that occurs when 272 the default Noah snow albedo scheme is used. Where the default Noah scheme results 273 in a warm bias, however, the improved albedo scheme does not improve the accuracy 274 of the WRF estimates. Scatterplots comparing air temperature observations from CMA 275 276 station with WRF estimates made using the improved snow albedo scheme estimates, 277 show the data to be concentrated near the ideal fitting line, where the model has exactly 278 reproduced the observations. Using the improved snow albedo scheme results in a 279 marked reduction in the cold bias for the WRF model estimates for EXP1, EXP2, EXP4, 280 EXP5 and EXP6, and the greatest reduction in the cold bias, for all eight experiments, 281 occurs for EXP6 (Fig. 2).





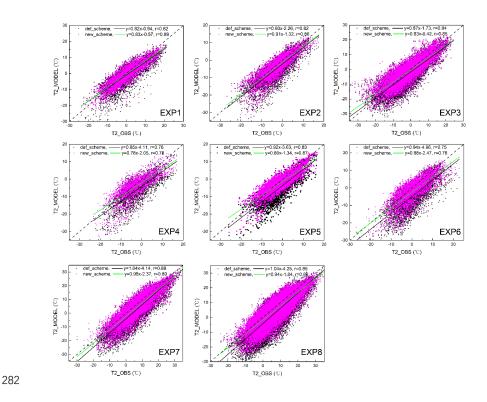


Figure 2. Scatterplot of air temperature, comparing the CMA observations and model 283 284 estimates for the eight experiments in the inner (high resolution) model domain from 285 WRF, using the default Noah snow albedo scheme (def_scheme, black solid circle), and using the improved snow albedo scheme (new_scheme, red solid circle). The black 286 287 solid line is a linear fit to the black solid circles. The green solid line is a linear fit to the red solid circles. The black dotted line is the line y=x. r is the correlation coefficient 288 between the CMA observations and the model estimates. The correlation coefficient (r) 289 is significant at 0.01 significance level. 290

To quantity the improvement to air temperature estimates that follows from implementation of the improved snow albedo scheme, the RMSE and correlation coefficient was calculated between CMA observations and model estimates of air temperature, shown in Figure 3. The differences between the accuracy of the new scheme and the default scheme are shown in Table 3. The accuracy of WRF air





296 temperature estimates varies between the different snowfall events, between the 297 different snow albedo schemes, and also varies with model resolution. The lowest air temperature RMSE and the highest correlation coefficient are 3.1 °C and 0.89, 298 respectively, and both occur for EXP1. The highest air temperature RMSE and the 299 lowest correlation coefficient occur for EXP6, reaching 7.2 °C and 0.75, respectively. 300 Air temperature RMSE generally ranges from 4.1 to 7.2 °C for the WRF estimates that 301 were made using the default Noah snow albedo scheme estimates, with correlation 302 coefficients ranging from 0.75 to 0.88. In contrast, when the improved snow albedo 303 scheme is implemented in WRF, the RMSE ranges from 3.1 to 5.6 °C, and the 304 correlation coefficients range from 0.78 to 0.89. Compared with when the default Noah 305 snow albedo scheme is used, the maximum decrease in air temperature RMSE when 306 the new scheme is used reaches 2 °C, which represents an improvement of 39 %, and 307 the average decrease in air temperature RMSE is 1.2 °C, representing an improvement 308 309 of 20.7 %. There is an improvement of more than 11 % in the RMSE for all eight 310 experiments when the new albedo scheme is used, relative to when the default scheme is used. Implementing the improved snow albedo scheme in WRF for all eight 311 312 experiments also increased the correlation coefficient between observed and modeled 313 air temperature, by 0.01-0.07, which represents an improvement of 1-9 % (Fig. 3, Table 314 3).

Compared with using the default Noah snow albedo scheme, using the improved 315 scheme results in improved model estimates for all eight EXPs, decreasing the air 316 temperature RMSE and increasing the correlation coefficient when compared with 317 318 observations. These improvements occur for air temperature estimates calculated at 319 both 5 km and 1 km resolution. The improvement to WRF model estimates is greater for calculations made at 1 km resolution than at 5 km resolution, and air temperature 320 321 estimates are more accurate at 1 km resolution than at 5 km resolution, regardless of 322 which albedo scheme is implemented (Fig. 3, Table 3). Therefore, fine resolution (i.e., 323 1 km) is strongly recommended for future snowfall event modeling studies.





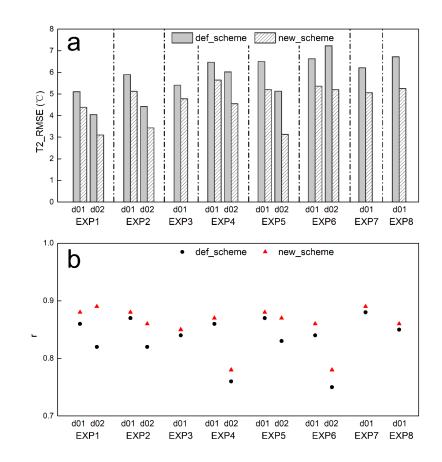




Figure 3. RMSE (a) and correlation coefficient (b) for air temperature (T2) between observations and model estimates in d01 and d02. The correlation coefficient is significant at the 0.01 significance level. For def_scheme and new_scheme, see Figure 2.

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333	Table 3 RMSE and correlation coefficient (r) for air temperature between CMA
334	observations and model estimates, calculated using the default Noah snow albedo
335	scheme (def_scheme) and the improved albedo scheme (new_scheme). The difference
336	in RMSE is new_scheme RMSE minus def_scheme RMSE. The difference in r is
337	new_scheme r minus def_scheme r. P value <0.05 is the significance test for the
338	correlation.

	Model Domain	RMSE (°C)		Relative	r				
EXPs		def_	new_	D:ff	Difference	def_	new_	D:00	Relative
		scheme	scheme	Difference	of RMSE	scheme	scheme	Difference	Difference
	d01	5.11	4.38	-0.73	-14.3 %	0.86	0.88	0.02	2.3 %
EXP1	d02	4.05	3.1	-0.95	-23.5 %	0.82	0.89	0.07	8.5 %
EXP2	d01	5.9	5.12	-0.78	-13.2 %	0.87	0.88	0.01	1.1 %
	d02	4.42	3.43	-0.99	-22.4 %	0.82	0.86	0.04	4.9 %
EXP3	d01	5.41	4.78	-0.63	-11.6 %	0.84	0.85	0.01	1.2 %
EXP4	d01	6.47	5.64	-0.83	-12.8 %	0.86	0.87	0.01	1.2 %
	d02	6.02	4.55	-1.47	-24.4 %	0.76	0.78	0.02	2.6 %
EVDC	d01	6.5	5.2	-1.30	-20.0 %	0.87	0.88	0.01	1.1 %
EXP5	d02	5.13	3.13	-2.00	-39.0 %	0.83	0.87	0.04	4.8 %
EVDC	d01	6.63	5.36	-1.27	-19.2 %	0.84	0.86	0.02	2.4 %
EXP6	d02	7.23	5.2	-2.03	-28.1 %	0.75	0.78	0.03	4.0 %
EXP7	d01	6.21	5.06	-1.15	-18.5 %	0.88	0.89	0.01	1.1 %
EXP8	d01	6.72	5.25	-1.47	-21.9 %	0.85	0.86	0.01	1.2 %

339 3.2 Albedo

Using the improved albedo scheme in the WRF model greatly reduces the cold air temperature bias that otherwise occurs, indicating an improvement to model performance. It is necessary to compare albedo estimates with in situ observations. There are very few observation stations located in the finer model domain, and so a total of 26 stations in d01 were used to evaluate the performance of the WRF



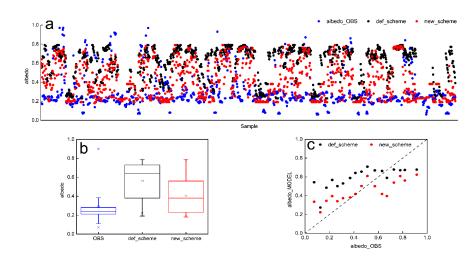


345 simulations of albedo at 5 km resolution (Table 2).

Scatterplots comparing observations and WRF estimates for albedo in the eight 346 experiments, when both the default and improved snow albedo scheme were used, are 347 348 shown alongside our statistical analysis in Figure 4. Albedo higher than 0.7 is interpreted as snowfall. Albedo in the range of 0.4 to 0.6 is interpreted as snowmelt. 349 350 Albedo lower than 0.3 indicates sparse or patchy snow cover at the in situ stations. For all eight snowfall events, the observed albedo is concentrated at low values, with a 351 median of 0.24, while WRF estimated albedo using the default Noah snow albedo 352 scheme has higher values, with a median of 0.64. Compared with albedo estimates 353 354 calculated using the default scheme, WRF estimates made using the improved scheme result in a prolonged snowmelt period, which increases the number of snowmelt 355 samples and leads to a median albedo of 0.38, which is closer to that for the in situ 356 observations. The mean albedo estimated from WRF using the improved scheme is 0.4, 357 which is also closer to the observed mean of 0.3, than the mean of 0.6 calculated from 358 359 WRF using the default scheme (Fig. 4a, 4b). In general, the accuracy of the WRF estimates when the new scheme is used is closely related to the observed albedo. 360 Compared with the WRF estimates made using the default Noah scheme, the WRF 361 estimates made using the improved scheme greatly reduce the overestimation of albedo 362 when the observed values are below 0.6, but seem to increase the underestimation of 363 albedo when the observed values are higher than 0.6 (Fig. 4c). 364







365

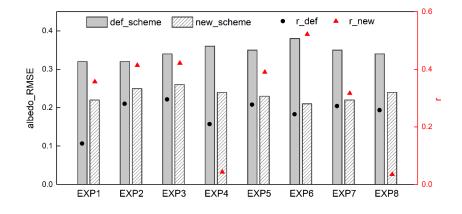
Figure 4. Scatterplot of albedo from observations and model estimates for the eight experiments (a: horizontal axis denotes samples from EXP1 to EXP8), corresponding box-and-whisker plot (b), and averaged observations and model estimates for the eight experiments (c). For def_scheme and new_scheme, see Figure 2.

370 To further evaluate the accuracy of WRF albedo estimates when the different snow 371 albedo schemes are used, we calculated the RMSE and correlation coefficients between the observations and the model estimates (Figure 5). The RMSE for the WRF estimates, 372 373 when compared to the observations, ranges from 0.32 to 0.38 for the eight experiments when the default scheme is used, and ranges from 0.21 to 0.26 when the improved 374 scheme is used. Compared with the default Noah snow albedo scheme, the improved 375 376 scheme results in a 0.1 decrease to the albedo RMSE in EXP1, representing a relative 377 decrease of 31.3 %; a 0.07 decrease in EXP2, representing a relative decrease of 21.9 %; a 0.08 decrease in EXP3, representing a relative 23.5 % decrease; a 0.12 decrease in 378 EXP4 and EXP5, representing relative decreases of 33.3 % and 34.3 %, respectively; a 379 0.17 decrease in EXP6, representing a relative 44.7 % decrease; a 0.13 decrease in 380 EXP7, representing a relative 37.1 % decrease; and a 0.1 decrease in EXP8, 381 representing a relative 29.4 % decrease. With the exceptions of EXP4 and EXP8, 382 correlations between the modeled and observed albedo are significant at the 0.01 383





384 significance level. Implementing the improved albedo scheme in WRF increases the albedo correlation coefficient by 0.21 in EXP1, a relative increase of 151 %; by 0.13 in 385 EXP2, a relative increase of 47.6 %; by 0.13 in EXP3, a relative increase of 42.5 %; by 386 0.11 in EXP5, a relative increase of 40.7 %; by 0.28 in EXP6, a relative increase of 387 114 %; and by 0.04 in EXP7, a relative increase of 16.2 % (Fig. 5). In general, during 388 snowfall and the snowmelt period that follows it, implementing WRF using the 389 improved snow albedo scheme outperforms implementing WRF using the default Noah 390 scheme and results in more accurate albedo estimates, demonstrated by considerable 391 decreases in RMSE and increases in the correlation coefficients. 392



393

Figure 5. RMSE in column and correlation coefficient (r) in scatterplot for albedo, comparing observations and model estimates. r_def and r_new mean r between the observations and WRF-estimates using the default albedo scheme, and using the improved albedo scheme, respectively. Except EXP4 and EXP8, the correlation coefficient is significant at the 0.01 significance level. For def_scheme and new_scheme, see Figure 2.

400 **3.3 Snow depth**

401 There is a feedback between albedo and snow. Snow accumulation and snowmelt 402 influence the proportion of solar irradiance that is reflected; albedo indirectly and non-

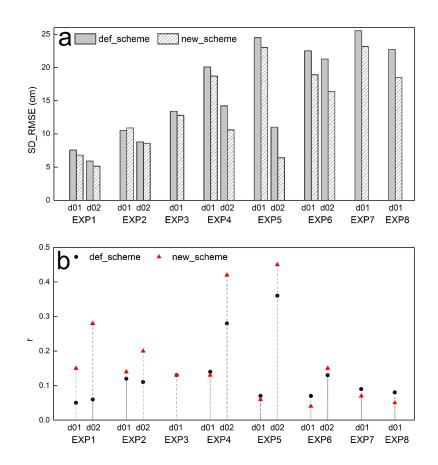




403 negligibly influences snow accumulation and snowmelt by affecting the land surface energy budget. In this study, we use an improved snow albedo scheme in which albedo 404 is parameterized as a function of snow depth and age. WRF model estimates of albedo 405 calculated using the improved snow albedo scheme outperform those calculated using 406 the default Noah scheme when snowfall events are simulated, and this leads to 407 improved representation of snowfall and the subsequent snowmelt processes in WRF 408 409 when the improved scheme is used. Instantaneous direct measurements of snow depth are recorded during snowfall events and over the subsequent snowmelt period. We use 410 these to quantity the improvement that using the new albedo scheme makes to snow 411 estimates calculated in WRF. We assess this by calculating the RMSE and correlation 412 coefficient between the model snow depth estimates and CMA observations, as shown 413 in Figure 6. 414







415

Figure 6. Same as Figure 3, but for RMSE (a) and correlation coefficient (b) for snow
depth (SD). The correlation coefficient is significant at the 0.01 significance level,
except for d01 estimates in EXP6.

Comparing the accuracy of WRF snow depth estimates calculated using the new scheme, to the accuracy achieved using the default Noah scheme, the greatest relative decrease in RMSE is 41.7 %, which occurs for estimates made at finer resolution in EXP5. Replacing the default albedo scheme with the new scheme in WRF results in an average decrease in snow depth RMSE of 2.2 cm, which is a 13.4 % improvement. In areas covered by the higher resolution model domain, the average RMSE decrease is 2.8 cm, which is a 21.2 % improvement. This shows that the impact of replacing the





426 albedo scheme with an improved scheme is more significant for areas in the higher resolution d02 model domain than for areas in the coarser d01 model domain (Fig. 6a). 427 Using the improved albedo scheme in WRF increases the correlation coefficient 428 429 between observed and modeled snow depth in areas within the d02 model domain, but this increase is not consistent for areas in the d01 domain. The greatest increase, both 430 relative and absolute, in the snow depth correlation coefficient occurs in the finer 431 simulation domain in EXP1, where the correlation between observed and modeled 432 snow depth increases by 0.22, which is a 366.7 % increase. The mean and relative 433 increases in the correlation coefficient between observed and modeled snow depth for 434 areas in the d02 simulations are 0.14 and 107.8 %, respectively (Fig. 6b). WRF snow 435 depth estimates are more accurate at finer resolution (i.e., 1 km resolution) than in 436 coarser simulations (i.e., 5 km resolution), regardless of which albedo scheme is 437 implemented (Fig. 6). Implementing the improved albedo scheme in WRF improves 438 439 the agreement between model estimated and observed snow depth, relative to 440 implementing the default Noah albedo scheme, as seen by the decreased RMSE and the 441 increased correlation coefficient.

442 **3.4 Turbulent heat and vapor fluxes**

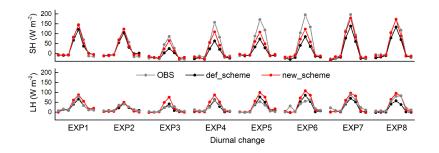
Albedo plays a significant role in the land surface energy balance. It determines the 443 proportioning of net radiation fluxes between turbulent heat and vapor fluxes. Our study 444 shows that using the improved snow albedo scheme in WRF results in a good model 445 representation of surface albedo in simulations of snow events. We now consider 446 whether replacing the default scheme with the improved scheme may affect the 447 proportioning of net radiation fluxes between turbulent heat and vapor fluxes. Since 448 there are very few observation in situ stations located in the area covered by the finer 449 model domain, a total of 14 in situ stations located in the area covered by d01 were 450 selected, and WRF estimates of turbulent heat and vapor fluxes were assessed from 451 simulations calculated at 5 km resolution (Table 2) 452

⁴⁵³ The diurnal changes in SH and LH recorded in the in situ observations and calculated





454 in the eight modeling experiments are shown in Figure 7. The WRF model successfully captures the diurnal changes in SH and LH, particularly in EXP1 and EXP2, where the 455 model estimates of SH and LH are almost equal to the in situ observations. In the 456 nighttime, the WRF model accurately estimates SH and LH in all eight experiments. 457 However, during the day WRF consistently underestimates SH in all experiments 458 except EXP1 and EXP2, and estimates LH with varying accuracy when the default 459 Noah albedo scheme is used. For example, when the default Noah scheme is used, WRF 460 accurately estimates LH in EXP3 and EXP4, but overestimates LH in EXP5 and EXP6 461 and underestimates LH in EXP7 and EXP8. Compared with WRF estimates calculated 462 using the default scheme, WRF simulations calculated using the new albedo scheme 463 result in increased estimates of the turbulent heat and vapor fluxes. This leads to SH 464 estimates that are closer to observations for experiments EXP3 to EXP8, a greatly 465 overestimated LH for experiments EXP3 to EXP6, and a slightly overestimated LH, 466 467 which is closer to observations for EXP7 and EXP8 (Fig. 7). In general, WRF estimates of SH are improved through the implementation of the new albedo scheme, relative to 468 the default scheme, although SH is underestimated during snowfall events. The impact 469 470 of the improved albedo scheme on LH estimates varies between the different snowfall 471 events, but LH is consistently overestimated.



472

Figure 7. Diurnal change of sensible (SH) and latent (LH) heat fluxes from ground
observations and model estimates. Beginning at 02:00 BST with 3 hours interval. For
def_scheme and new_scheme, see Figure 2.



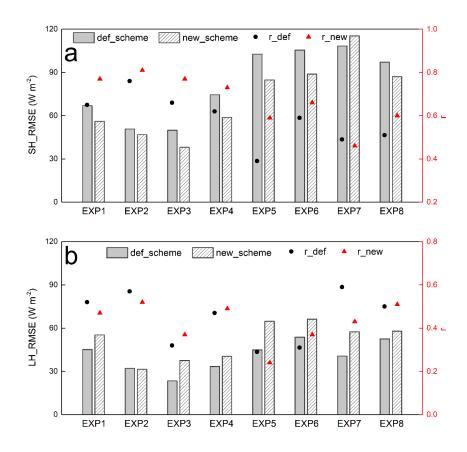


476	The RMSE and correlation coefficient for comparisons between WRF estimates and
477	observations for SH and LH are shown in Figure 8. Compared with when the default
478	Noah scheme is used, WRF estimates using the improved scheme result in reduced
479	RMSE for SH estimates in all experiments except EXP7. The absolute (and relative)
480	reductions are: 10.9 W m $^{-2}$ (16.2 %) in EXP1, 3.9 W m $^{-2}$ (7.6 %) in EXP2, 11.9 W m $^{-2}$
481	(23.8 %) in EXP3, 15.7 W m ⁻² (21 %) in EXP4, 17.9 W m ⁻² (17.4 %) in EXP5, 16.6 W
482	$m^{\text{-}2}$ (15.8 %) in EXP6, and 10.0 W $m^{\text{-}2}$ (10.3 %) in EXP8. There is a relative increase
483	of 6.5 % in EXP7. Implementing the improved scheme in WRF significantly increases
484	the correlation coefficients between observed and modeled SH, relative to using the
485	default scheme, in all experiments except EXP7, where there is a slight decrease. The
486	largest increase in the SH correlation coefficient, both absolute and relative, is 0.2 and
487	51.3 %, respectively, and occurs for EXP5. Implementing the improved scheme in WRF
488	reduces the SH RMSE by an average of 10 W m $^{\text{-2}}$, which is an improvement of 13.2 $\%$
489	improvement, and increases the SH correlation coefficient by an average of 0.1, which
490	is an improvement of 16.8 % (Fig. 8a). However, replacing the default scheme with the
491	improved albedo scheme results in less accurate estimates of LH, and corresponds to
492	an increase in RMSE in almost all eight experiments, although the correlation
493	coefficient increases for half of the snowfall events simulations (EXP3, EXP4, EXP6
494	and EXP8) (Fig. 8b).

In summary, the improved snow albedo scheme has a significant effect on the proportioning of radiative fluxes between turbulent heat and vapor fluxes. It significantly outperforms the default Noah scheme in relation to SH estimates, but there is no significant improvement in LH estimates and these may be less accurate when the new scheme is used, relative to the default scheme, during snowfall and the subsequent snowmelt period.







501

Figure 8. Same as Figure 5, but for sensible (a) and latent (b) heat fluxes. The
correlation coefficient is significant at the 0.001 significance level.

504 4 Discussion

The highly complex topography of the Tibetan Plateau means that WRF estimates of air temperature, albedo and snow depth are strongly sensitivity to model resolution. Our study shows that WRF performs much more accurately when run at finer resolution (1 km) than at relatively coarse resolution (5 km) for snowfall events over the Tibetan Plateau, regardless of which snow albedo parameterization scheme is used. This difference may be explained by the ability of the model to resolve complex terrain (Rahimi et al., 2019), and/or by the implementation of the cumulus convective





512 parameterization scheme. The more detailed representation of complex terrain and the explicit representation of convection mean that running WRF at finer resolution greatly 513 improves model estimates of air temperature, surface pressure, and relative humidity 514 515 (Singh et al., 2020), and provides small improvements in the magnitude of daytime convective precipitation (Collier and Immerzeel, 2015). Norris et al. (2017) pointed out 516 that decreasing the grid spacing from 6.7 to 2.2 km likely improves estimates of 517 mountain precipitation but does not fundamentally change the representation of the 518 diurnal cycle. They indicated that the key difference between low and high model 519 resolution is whether or not a cumulus convective scheme is required. Subkilometer 520 grid resolution has been investigated in WRF, and used for modeling meteorological 521 variables over complex terrain (Horvath et al., 2012; Dimitrova et al., 2016). The 500 522 m resolution configuration of WRF results in the closest match between the model 523 estimates and observations, and gives the most plausible spatial distribution of 524 525 precipitation over the complex topography. The performance of the WRF model has been similarly demonstrated to improve at 500 m, relative to coarser resolutions, for 526 wind and air temperature estimates (Bonekamp et al., 2018). We therefore strongly 527 528 suggest that subkilometer grid resolution should be considered when WRF is 529 configured for simulations covering areas in High Mountain Asia in future research.

530 Our improved snow albedo scheme parameterizes albedo as a function of snow depth and age by considering the relationship between MODIS albedo and the modelled snow 531 depth and age. It is more physically plausible than the default Noah scheme, which 532 considers snow cover, and outperforms the default Noah scheme for air temperature, 533 534 snow depth, albedo and turbulent heat and vapor fluxes estimates during snowfall 535 events. However, even when the improved albedo scheme is used, the RMSE for WRF estimates of albedo at 5 km spatial resolution remains around 0.21-0.26, although this 536 537 represents a decrease of 22-45 % relative to when the default albedo scheme is used. It should be noted that the accuracy of the MODIS albedo retrieval algorithm is limited 538 during snowfall events and snowmelt periods (Qin et al., 2011; An et al., 2020), and 539 540 also that rugged mountain terrain not only affects the radiation absorbed by the land





541 surface, but also affects the radiation reflected by the land surface to the satellite borne sensor. Multiple reflection and scattering from adjacent mountains creates challenges 542 for the monitoring and retrieval of surface albedo in areas of complex terrain via remote 543 sensing (Zhang and Gao, 2011; Roupioz et al., 2014, 2016). This reduces the accuracy 544 of MODIS albedo retrieval over the complex topographic Tibetan Plateau and 545 constitutes a limitation to the improved albedo parameterization investigated here, since 546 it relies on MODIS albedo products. A terrain correction is required for the MODIS 547 albedo retrieval to further improve the albedo parameterization scheme used here. 548 However, it is difficult to establish a unified terrain correction model due to the large 549 undulations of the Tibetan Plateau. How to effectively eliminate the influence of terrain 550 factors from a specific mountain surface on quantitative retrievals from remote sensing 551 data has long been a challenge and a focus for remote sensing research. A further 552 challenge for the assessment presented here is the sparse and uneven distribution of 553 554 available in situ albedo observation data over the bulk Tibetan Plateau. This paucity of 555 data means that there is a mismatch in spatial resolution when comparing albedo 556 estimates, calculated at 5 km resolution, with in situ observations.

Air temperature is a critical factor that is related to albedo and determines energy 557 distribution between SH and LH. Using the improved snow albedo parameterization 558 559 scheme significantly reduces the albedo overestimates during snowfall events that occur when the default scheme is used, and this leads to the reduction of the cold air 560 temperature bias in the model. Therefore, in this study, the improved scheme reduces 561 the underestimates for SH and improves the performance of WRF for simulating SH 562 563 over the Tibetan Plateau, relative to when the default scheme is used. These results 564 indicate that the accurate simulation of surface albedo is very important for the accurate 565 simulation of SH.

566 Implementing the improved albedo scheme results in little improvement to estimates of 567 LH calculated, which is restricted by water content and increased only slightly, relative 568 to when the default albedo scheme is used. This may be explained by the LH





569 parameterization scheme used in the Noah LSM (Chen and Dudhia, 2001). The total 570 LH in the Noah LSM has three components (LH from the direct evaporation from the top surface layer, evaporation of precipitation intercepted by the canopy and 571 572 transpiration via the canopy and roots respectively). The factors affecting calculation of LH in the Noah LSM include not only the radiation balance (which is impacted by 573 574 albedo), but also soil water, soil capillary conductivity and vegetation status, i.e., albedo, surface heat and water vapor exchange coefficient, saturated water vapor pressure, 575 specific humidity, surface soil water content, field capacity, wilting point, canopy 576 resistance, total precipitation, and canopy interception amount. In our current study, we 577 have focused on the snow albedo parameterization scheme in the Noah LSM by 578 considering the MODIS albedo product and the additional snow related variable of 579 snow depth. Therefore, the influence of our improved scheme on LH estimates 580 calculated by the LSM is very limited. 581

Implementing the improved snow albedo scheme in place of the default scheme greatly decreases the overestimation of albedo from snowmelt to snow free processes, but does not remove the underestimation of albedo during snowfall. This means that the improvements mainly come from snowmelt and snow free simulations, and model performance during snowfall may be worse when the improved albedo scheme is used. This suggests an opportunity to further investigate how albedo is characterized by snow depth and age in the snow albedo parameterization scheme.

589 **5 Conclusions**

We conducted several numerical experiments to evaluate the performance of the MODIS albedo based snow albedo parameterization scheme (Liu, 2020) implemented in WRF. We assessed the RMSE and correlation coefficient between observed and modeled air temperature, albedo, snow depth and turbulent heat and vapor fluxes for simulations of eight snowfall events over the Tibetan Plateau. We compared the accuracy of WRF estimates made using the improved snow albedo scheme with that of WRF estimates made using the default Noah scheme, in both cases comparing with





597 ground observations.

598 The accuracy of WRF estimates of albedo is significantly improved when the new albedo scheme is implemented. The default Noah scheme tends towards higher albedo 599 600 estimates and cannot accurately capture snowfall and snowmelt processes, resulting in a high RMSE and low correlation coefficient between modeled and observed albedo. 601 602 Through consideration of snow related variables, such as snow depth and age, and by being based on MODIS remote sensing albedo products, the improved scheme 603 estimates albedo more accurately than the default scheme, improving the albedo RMSE 604 by around 22 % to 45 %, with an average improvement of 32 %. Similarly, the improved 605 606 scheme results in an increased correlation coefficient between modeled and observed albedo. Relative to the default scheme, the correlation coefficient relatively increases 607 by around 16 % to 151 %, with an average improvement of 69 %. This may contribute 608 to the relatively better performance of WRF for simulating air temperatures when the 609 improved albedo scheme is used. The improved scheme relatively decreases (increases) 610 the air temperature RMSE (correlation coefficient) by 16 % (1.5 %) for model estimates 611 calculated at 5 km resolution, and by 27 % (5 %) for model estimates calculated at 1 612 km resolution. 613

There are mutual feedbacks between snow and albedo. During snowfall and over the 614 subsequent snowmelt period, snow depth and age affect changes in the albedo. The 615 changes in albedo in turn affect the evolution of snow events by changing the surface 616 energy budget and the proportional distribution of net radiation between turbulent heat 617 and vapor fluxes, and finally by changing the type of precipitation. Our study shows 618 that when the default albedo scheme is replaced by the improved albedo scheme in 619 WRF, the turbulent heat and vapor fluxes estimates increase. The improved scheme 620 significantly outperforms the default Noah scheme for SH estimates, with a reduction 621 (increase) in the RMSE (correlation coefficient) of 10 W m⁻² (0.1), representing an 622 improvement of 13 % (17 %). The overall accuracy with which WRF estimates 623 turbulent heat and vapor fluxes improves when the improved albedo scheme replaces 624





625 the default scheme, although there is no significant improvement in LH estimates. This overall improvement leads to a more accurate reproduction of the evolution of snowfall 626 events and to more accurate snow depth estimates. Our study shows that using the 627 628 improved albedo scheme in WRF reduces the RMSE and increases the correlation coefficient between modeled and observed snow depth, relative to using the default 629 scheme. This improvement is more significant for simulations at 1 km resolution than 630 for simulations at 5 km resolution, with maximum and averaged relative RMSE 631 (correlation coefficient) decreases (increases) of 42 % (367 %) and 21 % (108 %), 632 633 respectively for 1 km resolution simulations.

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653 paper.

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