

1 Topographic effects on solar radiation distribution in 2 mountainous watersheds and their influence on reference 3 evapotranspiration estimates at watershed scale

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9 10 **Abstract**

11 Distributed energy and water balance models require time-series surfaces of the
12 climatological variables involved in hydrological processes. Among them, solar radiation
13 plays an important role, especially in arid environments, as it is a key variable to the
14 circulation of water in the atmosphere. The lack of reliable data for the assessment of solar
15 radiation has led to the use of models. Most of the hydrological GIS-based models apply
16 simple interpolation techniques to data measured at sparse meteorological stations
17 disregarding topographic effects. Here, a topographic solar radiation algorithm is included for
18 the generation of detailed time-series solar radiation surfaces using limited data and relatively
19 simple methods, in order to quantify the effects of topography on the water losses through
20 evapotranspiration estimates in a mountainous watershed in southern Spain. First, the
21 comparison between the topographically corrected interpolated values of daily solar radiation
22 and those obtained by a direct spatial interpolation technique (Inverse Distance Weighed,
23 IDW) is provided. The results show the major role of topography in local values and
24 differences of up to +60% and -90% in the estimated daily values. Besides, the results are
25 compared to experimental data proving the usefulness of the model for the estimation of
26 spatially distributed radiation values in complex terrain, with a good adjustment for daily
27 values and the best fits under cloudless skies at hourly time steps. Finally, evapotranspiration
28 fields estimated through the ASCE-Penman-Monteith equation using both corrected and non-
29 corrected radiation values address the hydrologic importance of using topographically

1 corrected solar radiation fields as inputs to the equation over uniform values with mean
2 differences in the watershed of 62 mm/year and 142 mm/year of standard deviation. High
3 speed computations in a 1300 km² watershed in the south of Spain with up to a one-hour time
4 scale in 30 x 30 m² cells can be easily carried out on a desktop PC.

5

6 **1 Introduction**

7 There are several methods available for the development of digital elevation models for
8 hydrological studies but regular grid structures provide the best compromise between
9 accuracy and computational efficiency (Moore et al., 1991). For this, all the inputs to
10 distributed hydrological modelling must be available at this spatial scale. Among such inputs
11 to hydrological models, solar radiation plays an important role in most of the processes
12 involved, as it is a key variable in the circulation of water from the earth's surface to the
13 atmosphere, especially at Mediterranean regions. At a global scale, latitudinal gradients
14 caused by the earth's rotation and translation movements are well-known. However, at a
15 smaller scale, apart from cloudiness and other atmospheric heterogeneities, topography
16 determines the distribution of the incoming solar radiation; variability in slope angle and
17 slope orientation, as well as the shadows cast by topographic agents, can lead to strong local
18 gradients in solar radiation (Dozier, 1980; Dubayah, 1992; Dubayah and van Katwijk, 1992),
19 with the corresponding influence on the energy-mass balance of the snow cover and its
20 evolution (Dubayah and van Katwijk, 1992; Herrero et al., 2009), the vegetation canopy
21 (Dubayah, 1994), the surface soil layer, surface water bodies, etc.

22 The regional climate in Mediterranean areas is characterized by great inland-coast, valley-hill
23 contrasts, and is subject to cyclical fluctuations in cloud cover, precipitation and drought, thus
24 exhibiting considerable spatial and temporal variations (Diodato and Bellocchi, 2007). In such
25 latitudes, during periods of lack of rainfall - a common event at different spatial and temporal
26 scales - radiation is the main force in the system which causes both snowmelt and
27 evapotranspiration. Here, an accurate estimation of time-series solar radiation surfaces is
28 required for distributed energy and water balance modelling (Ranzi and Rosso, 1995; Herrero
29 et al., 2007).

30 One of the main drawbacks in the assessment of solar radiation is the lack of reliable data. In
31 mountainous areas where the monitoring network ineffectively covers the complex
32 heterogeneity of the terrain, simple geostatistical methods for spatial interpolation are not

1 always representative enough, and algorithms that explicitly or implicitly account for the
2 features creating strong local gradients in the incoming radiation must be applied (Susong et
3 al., 1999; Garen and Marks, 2005; Chen et al., 2007). Thus, the implementation of the spatial
4 variability in the incoming radiation at the cell scale for distributed hydrological modelling is
5 of major concern, especially in mountainous areas (Allen et al., 2006). Here, the combination
6 of extreme gradients in the spatial distribution of solar radiation, together with the lack of
7 measurements at detailed spatial and temporal scales, calls for the integration of algorithms
8 simple enough to be run with common measurements but at the same time able to capture the
9 agents that constitute the main sources of the spatial and temporal variability of solar
10 radiation.

11 At the local scale, the amount of solar radiation reaching a given location is called global
12 solar radiation and it depends mainly on the cloud cover, the turbidity of the clean air, the
13 time of the year, latitude, and surface geometry (Iqbal, 1983; Essery and Marks, 2007). As
14 radiation penetrates the atmosphere, it is depleted by absorption and scattering. Not all of the
15 scattered radiation is lost, since part of it eventually arrives at the surface of the earth in the
16 form of diffuse radiation (Liu and Jordan, 1960). Global radiation is the sum of direct or beam
17 irradiance from the sun, diffuse irradiance from the sky, where a portion of the overlying
18 hemisphere may be obstructed by the terrain, and direct and diffuse irradiance reflected by
19 nearby terrain (Dubayah, 1994). Therefore, global radiation received on a surface with a
20 random slope and aspect is largely controlled by atmospheric and topographic conditions
21 (Flint and Childs, 1987; Tian et al., 2001; Diodato and Bellocchi, 2007). In very rough terrain,
22 some areas may not receive any direct radiation during the whole year - even if facing south -
23 because of high peaks surrounding them. Under such conditions, a GIS-based solar radiation
24 model that considers the impact of terrain shading should be applied (Allen et al., 2006).

25 In any case, for the estimation of radiation incident on tilted surfaces, the partition of global
26 horizontal radiation into its beam and diffuse components is of major concern, as the
27 topographic effects are different for each one and therefore have to be modelled separately
28 (Iqbal, 1980; Antonic, 1998; González-Dugo et al., 2003). Thus, diffuse radiation is affected
29 by the unobstructed portion of the overlying hemisphere while reflected radiation is affected
30 by terrain slope and the portion of the overlying hemisphere obstructed by terrain (Dubayah,
31 1994). As for the beam component, self-shadowing and shadows cast by surrounding terrain
32 have to be considered for each sun position in the sky during the day.

1 For the quantification of the diffuse component, many parameters related to the atmospheric
2 properties are required in order to express the scattering properties of the atmosphere.
3 However, these parameters are not easily available or their computation from common
4 measurements may be time-consuming (Dubayah and van Katwijk, 1992) and so simpler
5 procedures need to be applied, especially at watershed scale and rough terrain. Thus, in these
6 situations the basic procedure for the partition of global radiation into its components is the
7 calculation of correlations between the daily global radiation and its diffuse component from
8 measured values of both quantities, and then to apply such correlations at locations where
9 diffuse radiation data are not available (Iqbal, 1980). In the literature there are several reviews
10 about the different correlations available, depending on the averaging procedure and on the
11 time scale of the radiation data (e.g. hourly or daily) (Iqbal, 1980; Spencer, 1982; Kambezidis
12 et al., 1994; Jacovides et al., 1996). Liu and Jordan (1960) were the first authors to develop a
13 model for the estimation of diffuse radiation from global data, establishing the basis for later
14 empirical analysis of global radiation from daily data. Ruth and Chant (1976) obtained a very
15 similar figure and demonstrated a latitudinal dependence in the models. Other authors
16 developed hourly correlations (Orgill and Hollands, 1977; Bugler, 1977; Erbs et al., 1982). In
17 1979 Collares-Pereira and Rabl, maintaining the assumption of isotropic approximation for
18 the diffuse radiation previously proposed by Liu and Jordan, improved some aspects of the
19 model (correction for the shade ring effect and use of daily values for extraterrestrial radiation
20 instead of single monthly values) and defined the daily clearness index as the ratio of global
21 radiation to extraterrestrial radiation.

22 Despite the availability of topographically corrected models for the estimation of solar
23 radiation fields as Dozier and Frew (1980), Dubayah (1992, 1994), etc., these approaches are
24 not commonly included in GIS-based hydrological models. On the contrary, most GIS-based
25 hydrological models usually adopt simple approaches to estimate the incident radiation
26 throughout the watershed. In AnnAGNPS, a distributed-parameter, physically- based,
27 continuous-simulation, watershed-scale, nonpoint-source pollutant model (Cronshey and
28 Theurer, 1998), correction factors to take account of the effects of dust, water vapour, path
29 length, and reflection and rescattering are applied to the extraterrestrial radiation in order to
30 obtain the short wave radiation received at the ground surface. However these corrections are
31 simplified into two multiplicative factors, one reflecting the effects of the atmosphere as a
32 function of the elevation and another as the influence of clouds, which depends on the
33 percentage of possible sunshine for each day (Bingner and Theurer, 2003). Land area

1 representations of a watershed are used to provide landscape spatial variability, so climatic
2 variables remain constant at a subwatershed scale and therefore do not involve topographic
3 factors. SWAT (Arnold et al., 1998) is a lumped model in which each subwatershed is
4 associated to a unique radiation gauge. Here, topographic corrections are not considered and
5 the measured solar radiation data, when available, are directly applied on the whole region of
6 influence by means of estimated extraterrestrial radiation. MIKE-SHE (Refsgaard and Storm,
7 1995), is a comprehensive, deterministic, distributed and physically-based modelling system
8 capable of simulating all major hydrological processes in the land phase of the hydrological
9 cycle (Singh et al., 1999). However, up to now, atmospheric processes have not generally
10 been modelled explicitly and whereas precipitation is a direct input in MIKE SHE, radiation
11 and water vapour transport in the atmosphere are typically bound up in evapotranspiration
12 models (Graham and Butts, 2005) and usually simple methods, such as Thiessen polygons or
13 other areal methods, are applied to extrapolate the point scale values for the referred stations
14 at a watershed scale (Singh et al., 1999; Vázquez et al., 2002; Vázquez and Feyen, 2003).

15 The aim of this study is to address the importance of incorporating in hydrological models
16 the effects of topography on the spatial distribution of global solar radiation at watershed
17 scale. To this purpose, different topographic algorithms have been coupled in order to
18 estimate series of distributed solar radiation values and calculations have been made to
19 quantify such influence on evapotranspiration estimates in mountainous areas in
20 Mediterranean locations. Thus, an algorithm was derived from Dozier (1980) and Jacovides et
21 al. (1996) to take into account the lack of meteorological stations at high altitudes. To be
22 exact, it should estimate hourly global values as well as the separation between its beam and
23 diffuse components from the common measurements obtained on horizontal surfaces. The
24 resulting algorithm was implemented on a GIS-based routine and applied to data from a
25 mountainous watershed on the south coast of Spain. The distributed results were compared to
26 those obtained from simpler interpolating methods and experimental data. Finally, in order to
27 address the hydrologic importance of using topographically corrected solar radiation fields
28 over uniform values, a simple evaluation in terms of their influence in the computation of
29 reference evapotranspiration fields is carried out.

30

1 **2 Materials and methods**

2 **2.1 Study area and data sources**

3 The study area is the Guadalfeo river watershed, Southern Spain (Fig. 1), where the highest
4 altitudes in Spain can be found (3 482 m) with the coastline only 40 km away, in a 1 300 km²
5 area which results in the interaction between semiarid Mediterranean and alpine climate
6 conditions, with the regular presence of snow (Díaz, 2007; Herrero, 2007; Aguilar, 2008;
7 Millares, 2008). The combination of such altitudinal gradients together with the large number
8 of vegetation, landforms and soil types produces a complex mountainous terrain with variable
9 hydrological behaviour. The main part of the watershed, in terms of hydrology, is comprised
10 of the southern hillside of Sierra Nevada, where global radiation is high throughout the year
11 due to its aspect and lack of cloud cover, even during winter, despite the cold temperatures
12 and the presence of snow. However, the deep valleys with a characteristic south-facing
13 orientation lead to important differences in the instantaneous global radiation between the
14 east- and west-facing mountain slopes, especially after sunrise and before sunset when these
15 valleys are mainly in the shade.

16 The meteorological data used in this study consisted of daily datasets provided by the three
17 stations (Fig. 1) of the Agroclimatic Information Network of Andalusia (RIA) available in the
18 watershed: 601, 602 and 603, whose UTM coordinates are shown in Table 1. Measurements
19 were made with a Skye Llandrindod Wells SP1110 pyranometer, with a characteristic range
20 of 0.35~1.1 μm .

21 The topographic input data are represented by a digital elevation model (DEM) with a
22 horizontal resolution of 30 x 30 m and 1 m of vertical precision (Fig. 1). Surface slope and
23 aspect were calculated for each point in the DEM, using the regression plane through the 3 x 3
24 neighbourhood of a given point after Dozier and Frew (1990).

25 For the evaluation of the algorithm performance, the daily datasets applied were provided by
26 one station of the Andalusian Alert and Phytosanitary Information Network (RAIF) (referred
27 as 702 in Fig. 1) which measures the variable with a Skye Llandrindod Wells SP1110
28 pyranometer, with a characteristic range of 0.35~1.1 μm , as well as hourly data recorded at a
29 new climatological station installed in 2004 in Sierra Nevada by the University of Granada
30 Environmental flow dynamics Research Group at an elevation of 2 510 m (referred as 802 in
31 Fig. 1). Measurements of global radiation at station 802 were made with a Kipp and Zonen

1 SP-Lite pyranometer, with a characteristic range of 0.4~1.1 μm . The sensor was placed on a
2 horizontal surface, partially surrounded by higher ground to the north but completely exposed
3 to the south.

4 **2.2 Calculation of solar radiation components**

5 The global radiation flux (R_g) aimed at a given location at a given moment is the sum of three
6 components: direct or beam radiation (R_b), diffuse radiation (R_d) and reflected radiation from
7 surrounding bodies (R_r).

8 The amount of radiation reaching the top of the atmosphere is inversely proportional to the
9 square of the distance to the sun, which is a fairly straightforward geometric procedure. The
10 solar constant (I_{CS}) is the rate at which solar energy affects a unit surface, at a normal angle to
11 the sun's rays, in free space, at the earth's mean distance from the sun. In fact, I_{CS} is not a
12 constant value but it can be fixed around $1\,367\text{ Wm}^{-2}$ (Frölich and Brusa, 1981) for practical
13 purposes. For the extraterrestrial radiation incident upon a horizontal unit surface normal to
14 the sun's rays (R_{on}), in other words, the radiation which would be incident on the same
15 horizontal surface in the absence of any atmosphere, I_{CS} is corrected with the eccentricity
16 factor (E_o) to account for the changes in distance from the earth to the sun along the elliptical
17 trajectory, according to the expressions in Iqbal (1983).

$$18 \quad R_{on} = E_o \cdot I_{CS} \quad (1)$$

19 Finally, for the extraterrestrial radiation incident upon a horizontal surface located at an angle
20 relative to the sun's beams (R_o), the solar coordinates (Fig. 2), the zenithal angle (θ_z) or its
21 complementary angle, the sun elevation angle (h_s), and the solar azimuth (ψ) have to be
22 previously defined:

$$23 \quad R_o = R_{on} \cdot \sin(h_s) = R_{on} \cdot \cos \theta_z \quad (2)$$

24 Solar coordinates, which are calculated following the equations in Iqbal (1983), are functions
25 of geographical latitude (ϕ_L), and solar declination (δ). Solar declination, or the angle
26 between the line joining the centers of the earth and the sun and the plane of the Equator,
27 varies from -23.5° to $+23.5^\circ$ during the year; however, as the maximum variation in one day is
28 never higher than 0.5° (at the equinoxes), a constant daily value can be used, which is the
29 value at the noon (Iqbal, 1983).

1 In order to obtain the total amount of global radiation during a day, extraterrestrial radiation
 2 must be integrated from sunrise to sunset. Daily extraterrestrial radiation values depend on the
 3 terrestrial coordinates but, as a medium-sized watershed, unique values for latitude and
 4 longitude were considered and therefore a constant daily value of extraterrestrial solar
 5 radiation was obtained for the whole watershed.

6 **2.2.1 Beam and diffuse component estimation on horizontal surfaces**

7 The correlation applied in the present study, shown in Eq. (3) and based on the clearness
 8 index (CI), was obtained by Jacovides et al. (1996), who investigated the accuracy of some of
 9 the previously available correlations when applied locally for high-quality data registered in
 10 Cyprus, and found that such correlations are location-independent. However, they developed
 11 a specific correlation which is more suitable for applying in Mediterranean areas:

$$12 \quad R_d/R_g = \begin{cases} 0.992 - 0.0486CI & CI \leq 0.1 \\ 0.954 + 0.734CI - 3.806CI^2 + 1.703CI^3 & 0.1 \leq CI \leq 0.71 \\ 0.165 & CI \geq 0.71 \end{cases} \quad (3)$$

13 Beam daily solar radiation (R_b) can be obtained as the difference between global and diffuse
 14 radiation:

$$15 \quad R_b = R_g - R_d \quad (4)$$

16 The application of hourly relations between hourly CI and hourly diffuse radiation values was
 17 initially considered following previous work in the literature (Orgill and Hollands, 1977;
 18 Bugler, 1977; Erbs et al., 1982). However, the aim of this work was to provide a feasible
 19 method to include topographic effects on radiation at watershed scale, and the size and
 20 heterogeneity of the study site together with the lack of meteorological stations, which
 21 unfortunately are usual circumstances in many locations, make it unreasonable to spatially
 22 interpolate hourly CI values as the spatial distribution of diffuse and direct radiation shows a
 23 better correlation at a daily scale. Besides, the application of hourly correlations would
 24 require the availability of some other variables such as solar altitude or air mass (González
 25 and Calbó, 1999). However, as pointed out by Zaksek et al. (2005), the use of more
 26 sophisticated models depends on the scale and purpose of the study, so that under certain
 27 circumstances it would be better to use a less complex model. Thus, a simpler approach is
 28 proposed so that once the daily values of each component are obtained for each cell, the
 29 hourly values (r_b and r_d), are computed by distributing the daily amounts along the day
 30 following the temporal pattern of extraterrestrial hourly radiation during the day. The hourly

1 values of beam and diffuse radiation on horizontal surfaces can then be transposed to give
 2 hourly radiation on tilted surfaces, since hourly methods of computing radiation on inclined
 3 planes, when available, should give slightly more accurate results than those obtained by the
 4 daily methods (Iqbal, 1978).

5

6 **2.2.2 Conversion from estimates on horizontal surfaces to tilted** 7 **surfaces**

8 With $\tau(\omega)$ a dimensionless transmission coefficient of the atmosphere for beam radiation on
 9 an horizontal surface, the following relationship between hourly beam values (r_b) and
 10 extraterrestrial radiation (r_o) for a certain hour angle (ω) both on an horizontal surface, is
 11 established (Iqbal, 1983):

$$12 \quad r_b = \tau_b(\omega)r_o \quad (5)$$

13 The same approximation, under the isotropic assumption on a randomly-oriented surface
 14 yields:

$$15 \quad r_{b,\beta\gamma} = \tau_b(\omega)r_{o,\beta\gamma} \quad (6)$$

16 with $r_{b,\beta\gamma}$ as the hourly beam radiation on a surface of slope β and orientation γ , and the same
 17 for the hourly extraterrestrial radiation ($r_{o,\beta\gamma}$) on a randomly-oriented surface. By replacing the
 18 transmission coefficient, the following expression for the estimation of $r_{b,\beta\gamma}$ in terms of the
 19 zenithal angle (θ_z) and a new corrected zenithal angle for the sloping surface (θ) is obtained:

$$20 \quad r_{b,\beta\gamma} = (r_b/r_o)r_{o,\beta\gamma} = r_b(\cos\theta/\cos\theta_z) \quad (7)$$

21 Therefore, for the calculation of hourly beam solar radiation on tilted surfaces, a correction in
 22 the solar coordinates is necessary, so that the cosine of the zenithal angle includes the effect of
 23 slope and orientation. Such corrected zenithal angle or illumination angle (θ), function of the
 24 sun-earth-tilted surface geometrical relationship, can be obtained as (Iqbal, 1983; Allen et al.,
 25 2006):

$$26 \quad \begin{aligned} \cos\theta = & \sin\delta \cdot (\sin\phi_L \cdot \cos\beta - \cos\phi_L \cdot \sin\beta \cdot \cos\gamma) + \\ & \cos\delta \cdot \cos\omega \cdot (\cos\phi_L \cdot \cos\beta - \sin\phi_L \cdot \sin\beta \cdot \cos\gamma) + \\ & \cos\delta \cdot \sin\beta \cdot \sin\gamma \cdot \sin\omega \end{aligned} \quad (8)$$

1 **2.3 Modelling topographic effects**

2 The topographic effects on solar radiation are mainly variations in the illumination angle and
3 shadowing from local horizons, the apparent intersection of the earth and the sky as seen by
4 an observer in a certain direction. The local horizon information from the gridded data allows
5 us to ascertain whether a given location at a certain sun position is shaded from direct sunlight
6 by surrounding terrain and determines, at any location, the portion of the overlying
7 hemisphere which is obscured by the terrain (Dozier et al., 1981; Dubayah, 1992). Thus, each
8 hourly component, beam, diffuse and reflected radiation is calculated separately to account
9 for the topographic effects (González-Dugo et al., 2003).

10 According to Essery and Marks (2007), even though since the availability of gridded data and
11 powerful computers many efficient algorithms for calculating distributions of solar radiation
12 over topographic grids have been developed, all of them implement the same basic geometric
13 principles. Thus, the calculation of horizons in this study was made following the
14 modification to the method by Dozier (1980), made more computationally efficient by Dozier
15 et al., (1981) and Dozier and Frew (1990). They developed a simple and fast algorithm for the
16 extraction of horizons from DEMs by comparing slopes between cells in a certain direction,
17 and formulated the problem by determining the coordinates of the points which constitute the
18 horizons in each cell. Then, by rotating the matrix and solving it in a one-dimensional way
19 along each row as many times as directions are considered, they derived the horizons in the
20 whole hemisphere for each cell in the DEM. In this study, eight directions were considered in
21 the calculation: the four cardinal points and their mid-way points.

22 **2.3.1 Beam radiation**

23 This fraction is strongly influenced by the illumination angle. Therefore, the main factors
24 conditioning the fraction of beam radiation are not only the slope and aspect of the location
25 relative to its neighbours, but also the location of the sun relative to the slope at each time
26 step. A certain location is receiving direct sunlight if none of the following situations are
27 taking place:

28 - Self-shadowing due to its own slope: this takes place if the vector normal to the surface
29 forms an angle greater than 90° with the solar vector (González-Dugo et al., 2003) (e.g. north-
30 facing hill slope and 45° slope, sun in the south at 30° over the horizon). This situation is easy
31 to calculate, as Eq. (8) yields a negative value.

1 - Shading cast by the nearby terrain: in this case, the sun is hidden by a local horizon. This
 2 case is more complex, since, unlike slope and orientation, it can not be calculated with
 3 information restricted to the immediate neighbourhood of a given point (Dozier et al., 1981).
 4 In order to express it mathematically, the term known as horizon angle in a certain
 5 direction ϕ , H_ϕ , is introduced as the angle between the normal to the surface and the line
 6 joining such point or grid in the DEM with another point in the same direction high enough to
 7 block solar radiation. Thus, shading by the surrounding terrain will occur for each time step if
 8 the illumination angle is greater than the horizon angle in that direction.

9 **2.3.2 Diffuse radiation**

10 Topography influences the diffuse component by modifying the portion of the overlying
 11 hemisphere visible at a certain point. The computation of scattered and reflected radiation
 12 fluxes from the atmosphere to the slopes is rather complicated, owing to the fact that the
 13 fluxes are considerably non-isotropic. A common assumption made is that the diffuse
 14 component of solar radiation (sky light) has an isotropic distribution over the hemispherical
 15 sky. However, the non-isotropic character of diffuse radiation fields (maximum intensities
 16 near the sun and the horizons, minimum intensities in the direction normal to that of the sun,
 17 etc) makes the simplified assumption sufficiently unrealistic to introduce errors into
 18 calculations of the energy incident on sloping surfaces (Temps and Coulson, 1977).
 19 Nevertheless, following the ideas of Kondratyev and Manolova (1960), who concluded that
 20 the isotropic approximation is sufficient for practical purposes (Klutcher, 1979), the isotropic
 21 assumption will prevail in this study with the portion of overlying hemisphere visible at each
 22 cell as the main factor controlling this component. Thus, the hourly diffuse radiation ($r_{d,\beta\gamma}$) on
 23 a surface of slope β and orientation γ , is:

$$24 \quad r_{d,\beta\gamma} = r_d \cdot SVF \quad (9)$$

25 where the sky view factor, SVF , is the ratio between the diffuse component at one given point
 26 and that on an unobstructed horizontal surface, so it corrects the incoming flux incident on a
 27 flat surface to flux over a sloping and possibly obstructed surface (Dubayah, 1992). Under the
 28 assumption of isotropic sky, a constant value for the SVF can be expressed analytically, in
 29 terms of the different horizons in each direction considered, as (Dozier and Frew, 1990):

$$30 \quad SVF = \sum_{\phi=1}^8 \cos \beta \cdot \sin^2 H_\phi + \left(\sin \beta \cdot \cos \gamma (H_\phi - \sin H_\phi \cos H_\phi) \right) \quad (10)$$

1 **2.3.3 Reflected radiation**

2 Albedo refers to the global reflectance of the surface to solar radiation. Both albedo and
3 topography can vary over short distances, and their interaction can lead to a wide variability
4 in global solar radiation on a scale of meters (Dubayah, 1992). Reflected radiation can be
5 computed following the ideas of Dozier and Frew (1990) from:

$$6 \quad r_{r,\beta\gamma} = \rho \cdot \left[\frac{(1 + \cos \beta)}{2} - SVF \right] \cdot (r_{d,\beta} + r_{b,\beta}) \quad (11)$$

7 where the term in brackets represents the terrain configuration factor for isotropic conditions
8 and infinitely long slope, and ρ is the albedo of the surface. The spatial average of albedo is a
9 factor which is difficult to estimate (Tasumi et al., 2006). In this study, the albedo was
10 estimated by Díaz et al. (2007) from the remote sensing data available from Landsat-5 and
11 Landsat-7 satellites during the study period. After the images have been properly corrected
12 and their reflectivity values extracted, albedo values are obtained at the cell scale through the
13 method proposed by Brest and Goward (1987) and interpolated for the whole time lapse on a
14 daily basis.

15 **2.3.4 Global radiation**

16 Finally, global radiation at an hourly scale is obtained as the sum of each component at an
17 hourly scale once: (1) direct irradiance has been corrected by self-shadowing and shadows
18 cast by nearby terrain; (2) diffuse sky irradiance has included the portion of the overlying
19 hemisphere that may be obstructed by nearby terrain, and (3) direct and diffuse irradiance
20 reflected by nearby terrain towards the location of interest have been calculated from both
21 corrected components (Dubayah, 1994).

22 **2.4 Evaluation of the topographic effects on solar radiation fields on** 23 **reference evapotranspiration estimates**

24 The choice of a method for the calculation of ET_0 depends on numerous factors. The available
25 energy at the soil surface is the first control of the process, so the estimation of this factor
26 from available data sometimes conditions the method (Shuttleworth, 1993). In this study, the
27 ASCE-Penman Monteith equation (Eq. (12)) was applied for the estimation of
28 evapotranspiration over a reference surface (Allen et al., 1998):

$$ET_0^{ASCE} = \frac{0.408\Delta(R_n - G) + \gamma \frac{C_n}{T + 273.16} u_2 (e_s - e_a)}{\Delta + \gamma(1 + C_d u_2)} \quad (12)$$

where ET_0^{ASCE} is the reference evapotranspiration during a certain time step (mm/ Δt); Δ the slope of the vapour pressure-temperature-curve saturation calculated at mean air temperature (kPa/ $^{\circ}$ C); γ the psychrometric constant (kPa/ $^{\circ}$ C); R_n and G the net radiation (combination of net shortwave and net longwave radiation) and soil heat fluxes, respectively, both in mm/ Δt water equivalent; e_a and e_s the actual and saturation vapour pressure (kPa), respectively; T the daily mean air temperature ($^{\circ}$ C) and u_2 the wind speed, both measured at a height of 2 metres above the soil surface (m/s). Finally C_d and C_n are resistance coefficients which vary with the reference crop, temporal time-step and, in the case of hourly time-steps, with daytime and night time. Here, the daily time step was applied and so the values of C_d and C_n were 900 and 0.34 respectively.

The calculation of some of the variables involved in the ASCE-PM equation can be found in detail depending on the available input data in Allen et al. (1998). Saxton (1975) found out that the variable to which the equation is most sensitive is net radiation. Net short-wave radiation (Eq. (13)) on the soil surface, as the difference between incident and reflected radiation, can be expressed in terms of the albedo of the surface, α (0.23 for the reference surface) and the predicted incoming solar global radiation, R_g (MJ/m²). In the same way, the net long-wave radiation was calculated by Eq. (14) where ϵ_{atm} and ϵ_{sup} are the atmospheric and surface emissivity respectively, T (K) the mean air temperature, T_{sup} (K) the temperature of the soil surface and σ Stefan-Boltzmann's constant ($4.903 \cdot 10^{-9}$ MJ/K⁴m²day). The atmospheric emissivity was calculated through a parametric expression by Herrero et al. (2009) based on near-surface measurements of solar radiation and relative humidity, valid for the local conditions of the study area. As ϵ_{sup} ranges from 0.985 in cotton crops to 0.94 in bare soil (Stefano and Ferro, 1997), a constant value of 1 was assumed. Besides, as soil temperature is not commonly available at broad scales it is assumed to be equal to the temperature of the air, and so the expression for net long-wave radiation remains as previously done by other authors (Doorenbos and Pruitt, 1977; Allen, 1986; Allen, 1998) as a modification to Stefan-Boltzmann's law due to the absorption and downward radiation from the sky. Thus, the product of the Stefan-Boltzmann's constant times the fourth power to the mean air temperature is modified with a cloudiness factor and an air humidity factor (Allen et al.,

1 1998; Donatelli et al., 2006), both factors included in this study in the term ε_{atm} and so,
2 together with the mean air temperature constitute the only inputs to the equation.

$$3 \quad R_{ns} = (1 - a)R_g \quad (13)$$

$$4 \quad R_{nl} = \varepsilon_{atm} \cdot \sigma \cdot T^4 - \varepsilon_{sup} \cdot \sigma \cdot T_{sup}^4 \approx (\varepsilon_{atm} - 1) \cdot \sigma \cdot T^4 \quad (14)$$

5 Finally, as the algorithm is able to derive global radiation values at the cell scale and once the
6 rest of inputs to the ASCE-PM equation are also available at the cell scale (Herrero et al.,
7 2007) the influence of topographic effects is evaluated in a distributed manner in ET_0^{ASCE} ,
8 estimated after using the topographically corrected values in comparison with distributed
9 estimates by IDW of the solar radiation data registered at the meteorological stations.

10

11 **3 Results and discussion**

12 In order to run the proposed set of algorithms at the watershed scale, hourly global radiation
13 was calculated from each 30 x 30 m² cell of the DEM in the study area for the period
14 comprised between 4 November 2004 and 29 April 2007.

15 The results are organised into three sections. Firstly, comparisons of the results obtained
16 through the topographic radiation algorithm previously exposed with those derived from a
17 classical interpolation technique are showed. Secondly, the suitability of the results at
18 different temporal scales is presented through its comparison with field measurements,
19 proving the accuracy of the estimated values for hydrological distributed modelling. Finally,
20 in order to address the hydrologic importance of using topographically corrected solar
21 radiation fields over uniform values obtained through classical interpolation techniques, the
22 influence of both estimations as inputs to reference evapotranspiration computations is
23 evaluated.

24 **3.1 Topographic corrections vs classic interpolation techniques on solar** 25 **radiation estimates**

26 At the first stage, topographic information was derived from the DEM (Fig. 1) in the study
27 area. Slope and orientation maps were obtained and the horizons for each cell were calculated
28 as stated before. Once these parameters are available for a certain area they can be used in
29 subsequent executions as they are considered to be independent of the time of the year.

1 In order to compare the results obtained through the topographic algorithm with those of
2 classic interpolation techniques, a reference day was selected. This date, 20/11/2004 was
3 chosen as it was cloudless and it had not rained for several days. This condition is very
4 important for the albedo estimation from remote sensing images, as the presence of moisture
5 in the environment influences the quality of the estimates and therefore consecutively dry
6 days are most suitable for an accurate performance. Moreover, remote sensing images were
7 available for this date and therefore the errors due to the temporal interpolation of albedo
8 values were minimized.

9 Combining the daily extraterrestrial value over a flat surface on the selected date (17.85
10 MJ/m²day) with the global radiation registered in the measuring network, which can be seen
11 in Table 1, the clearness index was obtained for each station and spatially interpolated
12 following the inverse distance weighed (IDW) method, in order to distribute it throughout the
13 watershed. This may appear to be quite an unrealistic simplification, but it is justified by the
14 lack of more spatially distributed registering sites which would allow to look into the factors
15 that affect the *CI* such as the variation in the atmospheric air mass through the height of each
16 cell, distance to the coast, etc. Therefore future research is proposed into the spatial
17 distribution of this index while a simple spatial interpolation technique is applied as a first
18 approximation in the present study.

19 Once the mean daily clearness index was spatially derived, global radiation values were
20 divided into its beam and diffuse components over a flat surface at a cell scale, and distributed
21 according to the hourly extraterrestrial radiation values for the subsequent topographic
22 corrections. Finally, the hourly sequence of global radiation, as the sum of each component at
23 an hourly scale once each component has been properly corrected, is shown in Fig. 3a, where
24 the spatial gradient in hourly global radiation is evident. On the whole, it can be seen that the
25 locations receiving more radiation are those in the highest part of the watershed, with a south-
26 facing orientation that remains unobstructed during most of the hours of daylight.

27 In order to assess the potential of the topographic corrections, a simple interpolation technique
28 was applied for the same date. For this, the inverse distance weighed was applied to the
29 hourly values of global radiation measured in the stations (Fig. 3b). From contrasting results
30 between Figs. 3a and 3b, not only was a huge difference visible in the distributed values of
31 the variable cell by cell, but also the wider range of global values in the watershed, when
32 topographic factors are taken into account. In this latter case, extreme values, far exceeding

1 the measured values, represent extreme conditions, such as high areas remaining unobstructed
2 most of the daytime and sometimes receiving almost double the values obtained through
3 interpolation of the data recorded at the stations or, at the other end of the scale, valleys that
4 receive minimal or even zero null quantities of solar radiation, due to the configuration of the
5 surrounding terrain. As a consequence, processes such as evaporation or snowmelt, which
6 rely heavily on solar radiation, can be miscalculated under a wide range of conditions, such as
7 overestimations in areas obstructed by nearby terrain or underestimations in the upper and
8 exposed regions of the watershed, among others.

9 Hourly values can be aggregated in each cell at the required temporal scale. In this way, Fig. 4
10 represents the spatial distribution of daily global radiation on 20/11/2004 estimated through
11 the topographic algorithm (Fig. 4a) and from IDW (Fig. 4b), respectively. Again, the same
12 ideas can be drawn as at an hourly time step, since maximum and minimum values found in
13 the watershed considering topographic effects are quite different to those obtained through
14 IDW and the daily values registered at the meteorological stations (Table 1). In this way, we
15 found differences of as much as an extra 60% in the estimated daily values compared with
16 those obtained through spatial interpolation without consideration of topography,
17 predominantly on the south-facing hillsides in the northern part of the watershed, and
18 estimates of up to 90 % of lower in certain cells obstructed most of the daytime.

19 Finally the same comparison for the hydrological year 2004, from 1 September 2004 to 31
20 August 2005 (Figs. 4c and 4d) resulted in a mean excess of 324 MJ/m²/year with a standard
21 deviation of 850 MJ/m²/year when applying IDW over the topographic computation and
22 extreme differences of the same order of magnitude as at the daily time step.

23 **3.2 Validation of topographic corrections**

24 The radiation values generated were compared against the radiation measurements and the
25 agreement between generated and measured data was evaluated through 1:1 lines. The period
26 considered in the evaluation was determined by the availability of data in the climatological
27 station 802, which included almost two and a half hydrological years, from 4 November 2004
28 to 29 April 2007.

29 As regards daily values, a close agreement between generated (R_{gp}) and measured (R_{go}) values
30 can be observed in Fig. 5 as points tend to line up around the 1:1 line. However, a slight
31 underestimation in the generated values can also be appreciated, with the topographic

1 approximation underestimating the values measured at stations 702 and 802 by 6 and 7%
2 respectively. Such underestimations take place especially in summer periods, as depicted in
3 Fig. 6, when the availability in this study of remote sensing images for an accurate estimation
4 of albedo was more limited, and also on very clear days, when an increase in global radiation
5 with altitude is expected, due to the reduction in density of the atmosphere, and when the
6 consideration of anisotropy in the atmosphere would modify the estimation of the diffuse
7 component. In this way, the consideration of factors creating spatial gradients in the
8 distribution of the *CI*, such as altitude, distance to the coast or proximity to urban areas, could
9 improve the results in our study, especially in very cloudless days and considering that the
10 climatological stations used to estimate the *CI* are located at relative low elevation compared
11 to the mean height of the watershed.

12 To conclude, the accuracy of predicted hourly values was assessed in station 802, where
13 measurements at this time scale were available. Despite the scattering effect observed in Fig.
14 7, which shows the agreement between predicted (r_{gp}) and measured (r_{go}) hourly values for
15 the evaluation period, we can say that the algorithm reasonably predicts the observed data
16 with a R^2 of 0.83, especially considering the time scale and some of the assumptions of the
17 algorithm which at this time step might appear rather simplistic. In this way, the installation
18 of a denser monitoring network provided with solar devices recording hourly direct and
19 diffuse radiation data may improve the results. Firstly, it would provide the spatial scheme
20 required for the spatial interpolation of hourly values. Secondly, it would allow including
21 more factors for the spatial distribution of the *CI* as previously suggested. Finally, the
22 derivation of hourly correlations between the hourly diffuse radiation and the *CI* would be
23 reasonable in the study area. To sum up, the possibility to work at finer scales would be the
24 ideal as the geometrical relationships involved in the calculation of extraterrestrial radiation
25 are continuous in time. However, independently of this continuous nature of extraterrestrial
26 radiation, the time scale of the computation of the incoming solar radiation is determined by
27 the temporal frequency of the monitoring network. Nevertheless, the calculations with
28 aggregated hourly values at higher temporal scales such as the daily time step showed the
29 same degree of detail that at hourly time scale (R^2 around 0.8).

30 Finally, since cloudless skies are required for an accurate characterization of the albedo from
31 remote sensing images, the results at an hourly time step were analysed considering this
32 effect. Thus, two different atmospheric situations in terms of the occurrence of rainfall are

1 defined as an indicator of the cloudiness in the atmosphere: events (when it rains somewhere
2 in the watershed) and non-events (periods between events). Fig. 8 represents hourly values for
3 event days on the left-hand side (a, b, c) and non-events on the right (d, e, f).

4 Table 2 shows different linear fits for each day represented in Fig. 8 and its calculated R^2
5 values. As was expected, the predicted values were much better for cloudless skies or non-
6 events, when acceptable R^2 values were obtained even when forcing the adjustment to reach
7 the origin. However, as with daily values, the algorithm slightly underestimated the observed
8 hourly values, which following the ideas of Ineichen and Pérez (2002) could be improved
9 with the consideration of the variation of atmospheric density with altitude, as the data used
10 for the calibration are registered at a climatological station located at a height of 2 510 meters.
11 In any case, the worst fits obtained in situations when events occur are expected to be more
12 closely related to the separation of the different components in the global radiation value than
13 to the topographic interpolation process.

14 Nevertheless, these results are considered to be acceptable in the framework of the present
15 study as the estimation of global radiation in semiarid environments is especially important
16 for cloudless days, when evaporative processes and snowmelt must be considered in water
17 and energy balance modelling. This is especially true considering that cloudless days
18 constitute a higher rate than cloudy days associated to situations when events occur in a
19 Mediterranean area like the present study site: in this case around 75% of clear sky days for
20 the evaluation period.

21 **3.3 Influence of the inclusion of topographic corrections on hydrological** 22 **variables: ET_0**

23 Finally, a distributed computation of ET_0 was applied for the same reference day of Sect. 3.1
24 (Fig. 9a) and the hydrological year 2004 (Fig. 9c) (1 September 2004 to 31 August 2005)
25 once all the variables involved in the ASCE-PM equation had been spatially derived including
26 topographic effects. Besides, the ASCE-PM equation was computed with solar radiation
27 surfaces obtained through IDW of the data registered at the stations as inputs to the equation
28 (Figs. 9b and 9d) in order to prove the importance of solar radiation fields which include
29 topographic corrections. Again, in Figs. 9a and 9c not only the apparent spatial variability of
30 ET_0 estimates cell by cell which follows the topographic gradient can be seen, but also a
31 wider range of values in the watershed than with IDW-interpolated solar radiation fields as

1 inputs (Figs. 9b and 9d). Besides, in this latter case, ET_0 estimates in the watershed appear to
2 be more influenced by the spatial distribution of other variables than by solar radiation (e.g.
3 temperature in Fig. 9b). Considering the mean statistics of the difference between both
4 computations on an annual basis, a mean excess of 62 mm/year and a standard deviation of
5 142 mm/year in ET_0 estimations when using IDW-interpolated solar radiation fields were
6 obtained. These differences in an area where the mean annual rainfall varies from 450
7 mm/year on the coast to 800 mm/year on the highest peaks may constitute a considerable
8 source of error in the water balance when applying distributed hydrological models for the
9 management and planning of water resources.

10

11 **4 Conclusions**

12 Difficulties are sometimes encountered in utilizing available solar radiation data, since they
13 consist primarily of total (direct plus diffuse) radiation only, and the knowledge of the values
14 for each component is often required, especially for the consideration of topographic effects
15 as they affect each component differently.

16 Thus, detailed time-series radiation surfaces have been developed, using limited data and
17 relatively simple methods, to drive distributed energy and water balance models in
18 mountainous Mediterranean environments. The interpolation is managed through linear
19 interpolation of CI as a clue to mean daily radiation, plus topographic properties
20 geometrically related to the sun's position at hourly intervals. Such calculations are easy to
21 reproduce from standard climatological station datasets. The significant incidence of
22 topography on the values of global solar radiation throughout the watershed has been
23 demonstrated by the results of the topographic solar radiation algorithm proposed. In this
24 way, differences of as much as an extra 60% in the estimated daily values compared with
25 those obtained through spatial interpolation without consideration of topography, and
26 estimates of up to 90 % of lower in certain cells obstructed most of the daytime were found.
27 This affects the modelling of the slow but extreme drying out of the watershed during periods
28 between events and the modelling of the snowmelt in the highest areas, among other
29 processes.

30 The simulated results fit well with the measured values of global radiation at the 2 510 meter
31 high monitoring point established for this work, with a correlation of 0.7 for daily values, and
32 an underestimation of 10% in days with extreme conditions, that decrease the validity of the

1 assumptions taken in the algorithm, as the previous paragraphs have justified. However, the
2 simulated results constitute a further approach to the accurate characterization of the spatial
3 distribution of hourly global radiation values in mountainous areas with scarce data
4 registering sites. On-going work will develop a further approach, and test the inclusion of
5 additional corrective terms through the establishment of two additional meteorological
6 stations equipped with pyranometers at points with increasing height above sea-level and
7 distance from the sea.

8 The importance of considering the topographic gradients in the spatial distribution of solar
9 radiation for the study of hydrological processes in which this variable plays a crucial role
10 became evident against ET_0 estimates with solar radiation fields obtained through classical
11 interpolation techniques of data registered at meteorological stations. In this way, a mean
12 excess of 62 mm/year was found with IDW-interpolated solar radiation fields as inputs to the
13 ASCE-PM equation.

14 Two final comments are included on the applicability of the algorithm proposed in this work.
15 For the purposes of computer simulation programs, which handle vast amounts of data, this
16 algorithm was implemented in Matlab during the trials and finally in C++ to get a sufficiently
17 fast computation, considering all the processes involved at a cell scale. Besides, some of the
18 assumptions that could appear quite unrealistic due mainly to the scarcity of data, have
19 managed to achieve a compromise between a sufficiently representative distributed
20 approximation and a high-speed processing algorithm that can be run on a desktop PC, from
21 the comparison with measured data and simpler interpolation techniques. Finally, through the
22 use of daily samples, the availability of data is enhanced as not many hourly registers are
23 needed; this allows the use of the algorithm in mountainous areas which lack a high frequency
24 monitoring network, which is so common in many other areas.

25

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28

1 Table 1. UTM coordinates of the climatological stations, measured daily global radiation and
2 clearness index on the 20/11/2004

3

Station	X	Y	Z	Rg (MJ/m ² day)	CI
601	483724	4086564	950	12.4	0.69
602	446712	4097327	781	14.4	0.80
603	439612	4066365	49	10.4	0.58
702	451435	4089276	700	13.46	0.75
802	471338	4098246	2510	13.81	0.77

4

1 Table 2. Linear fits of observed (r_{go}), and predicted (r_{gp}) global hourly radiation (MJ/m²) at
 2 station 802 for certain dates.

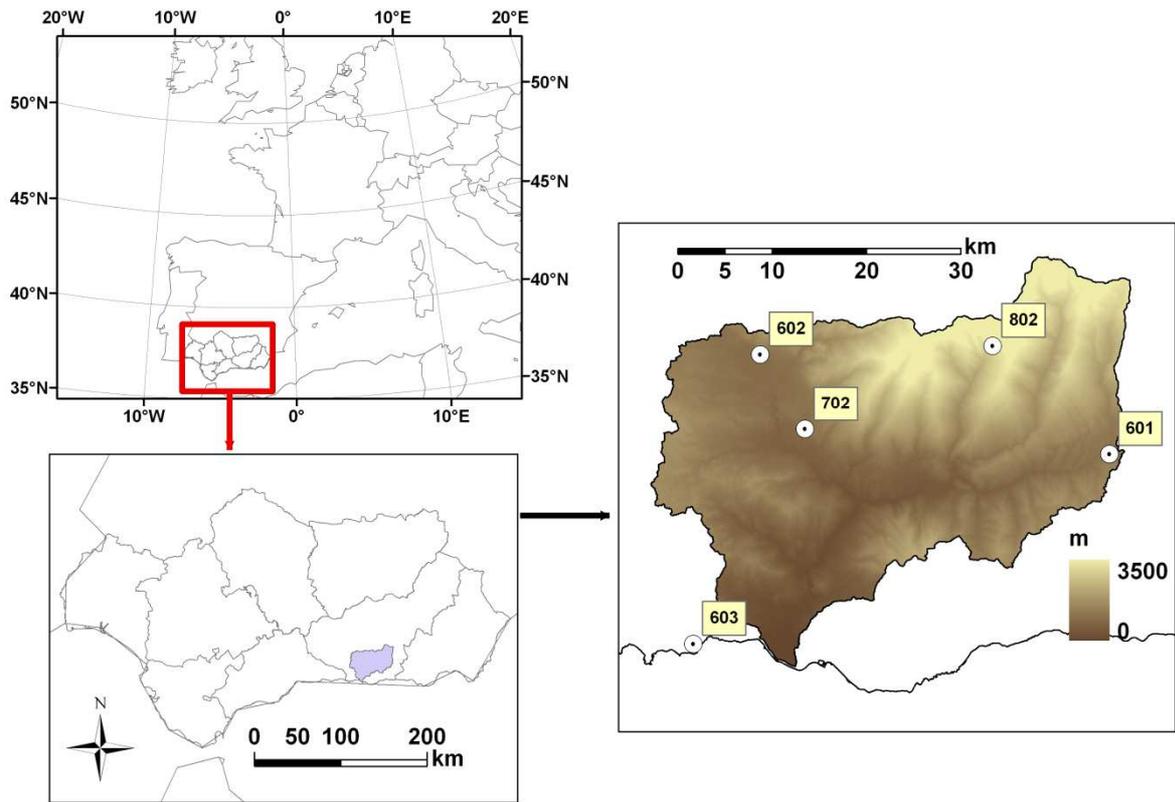
3

Equation type	$r_{gp}=a \cdot r_{go}$		$r_{gp}=a \cdot r_{go}+b$		
	a	R ²	a	b	R ²
Events					
a) 4/11/2004	1.26	0.69	1.69	0.34	0.75
b) 4/12/2004	0.40	0.17	0.23	0.26	0.62
c) 5/02/2005	0.88	0.19	0.31	0.71	0.21
Non events					
d) 15/11/2004	0.92	0.92	0.81	0.20	0.95
e) 15/12/2004	0.98	0.81	0.81	0.23	0.87
f) 20/02/2005	0.99	0.72	0.77	0.5	0.80

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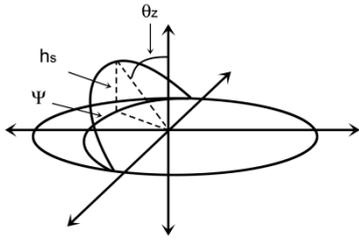


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4 Figure 1. Guadalfeo River Watershed, climatological stations and DEM

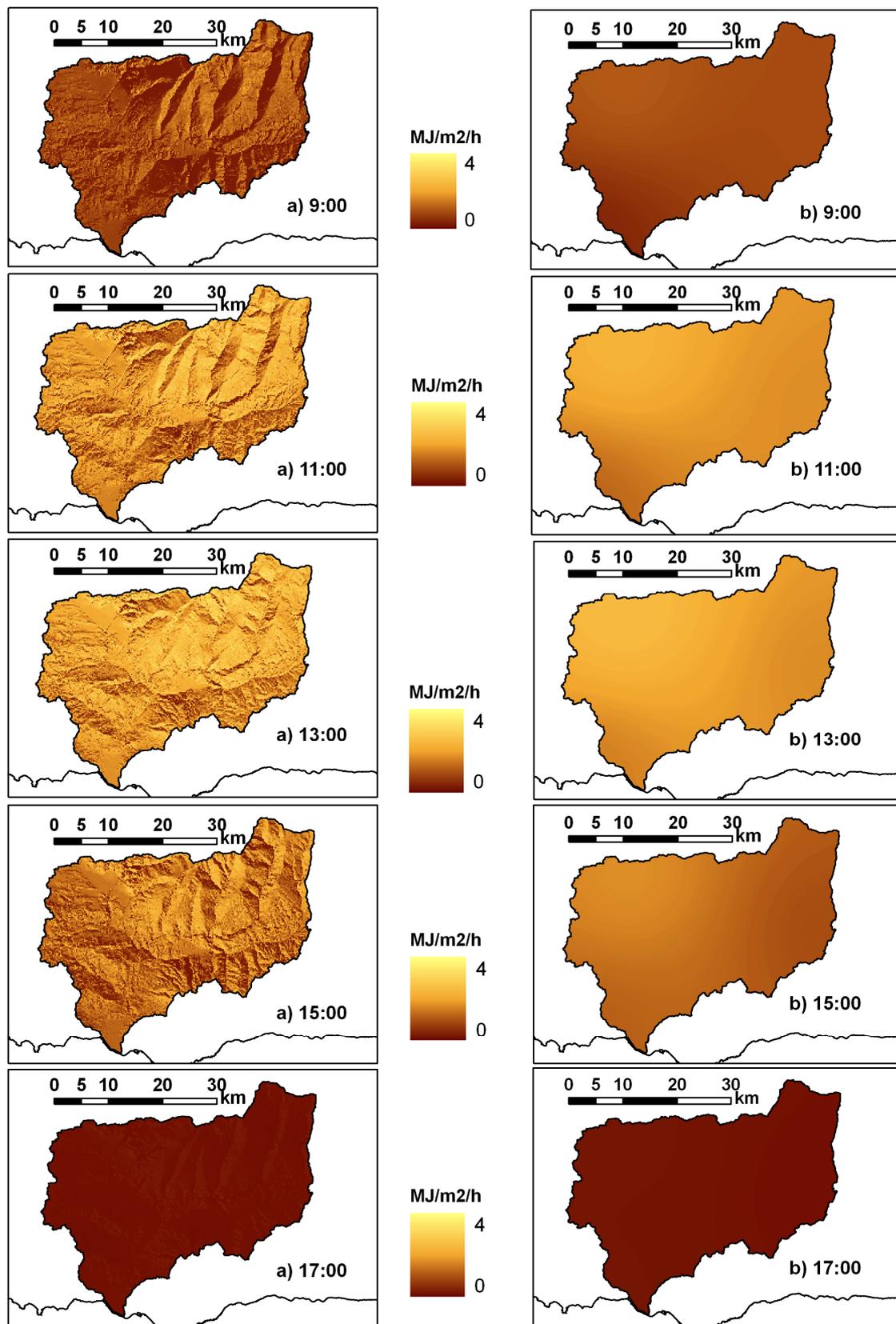
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2 Figure 2. Solar coordinates

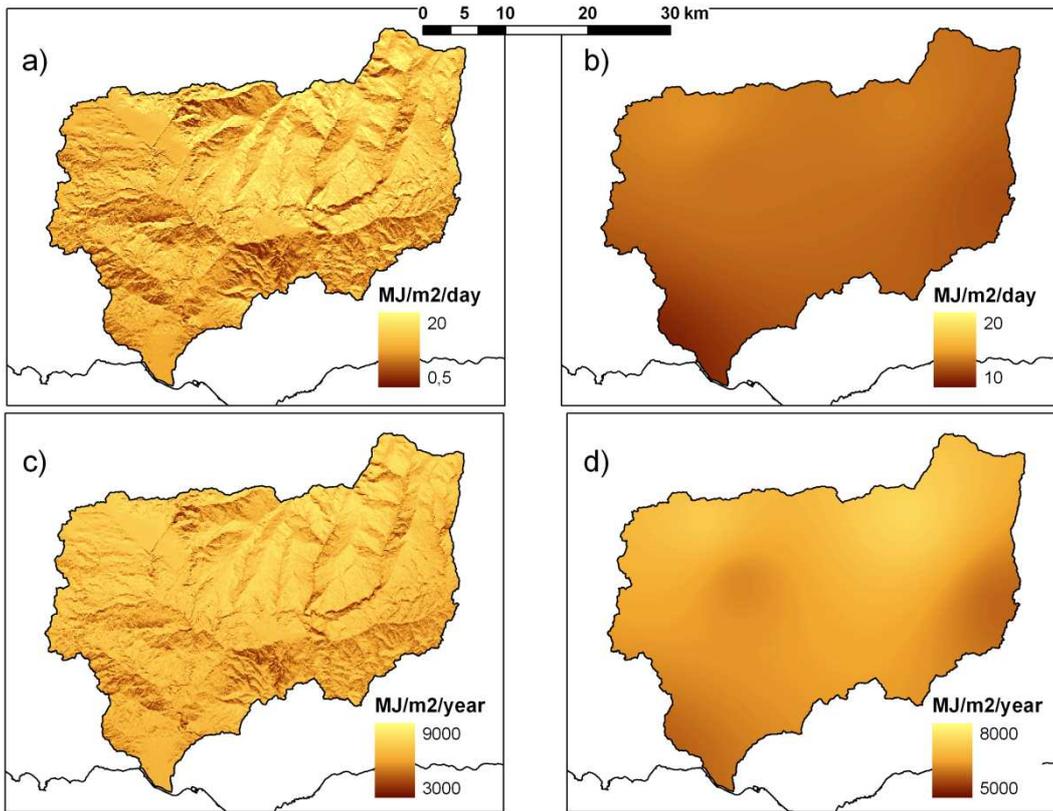
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2 Figure 3. Hourly global radiation (a) topographically corrected vs. (b) IDW interpolated
 3 (20/11/2004)

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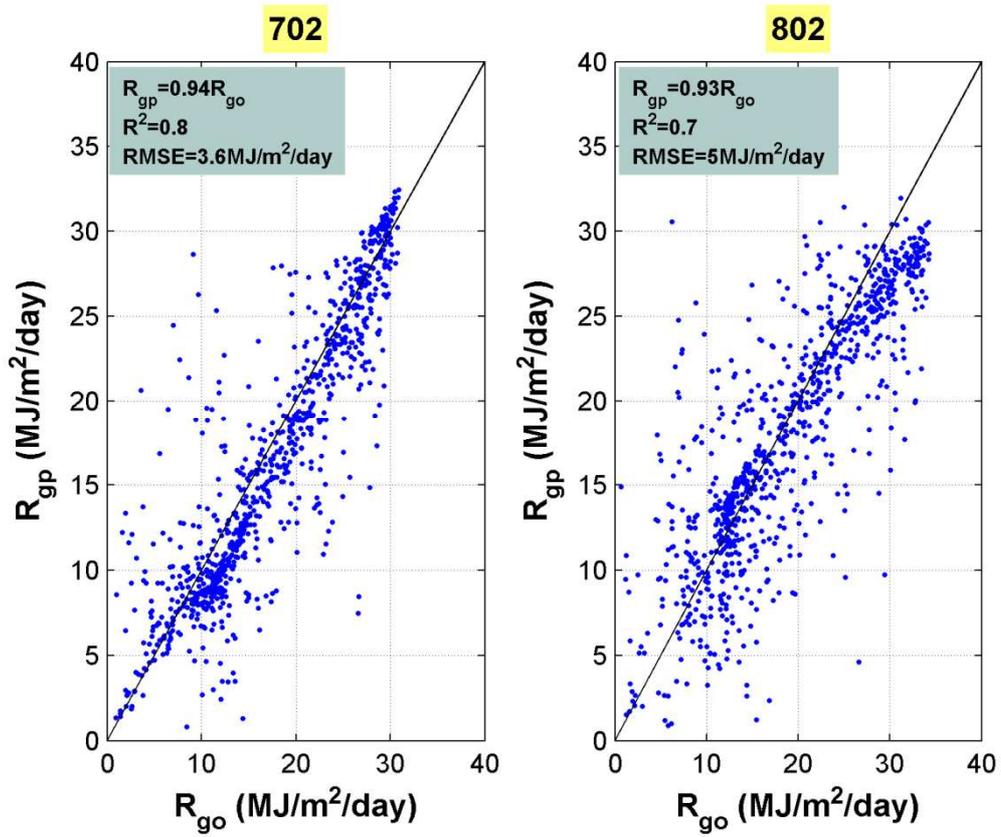
3 Figure 4. Daily global radiation (20/11/2004) a) topographically corrected vs. b) IDW
4 interpolated, and annual global radiation (1/09/2004-31/08/2005) (c) topographically
5 corrected vs.(d) IDW interpolated

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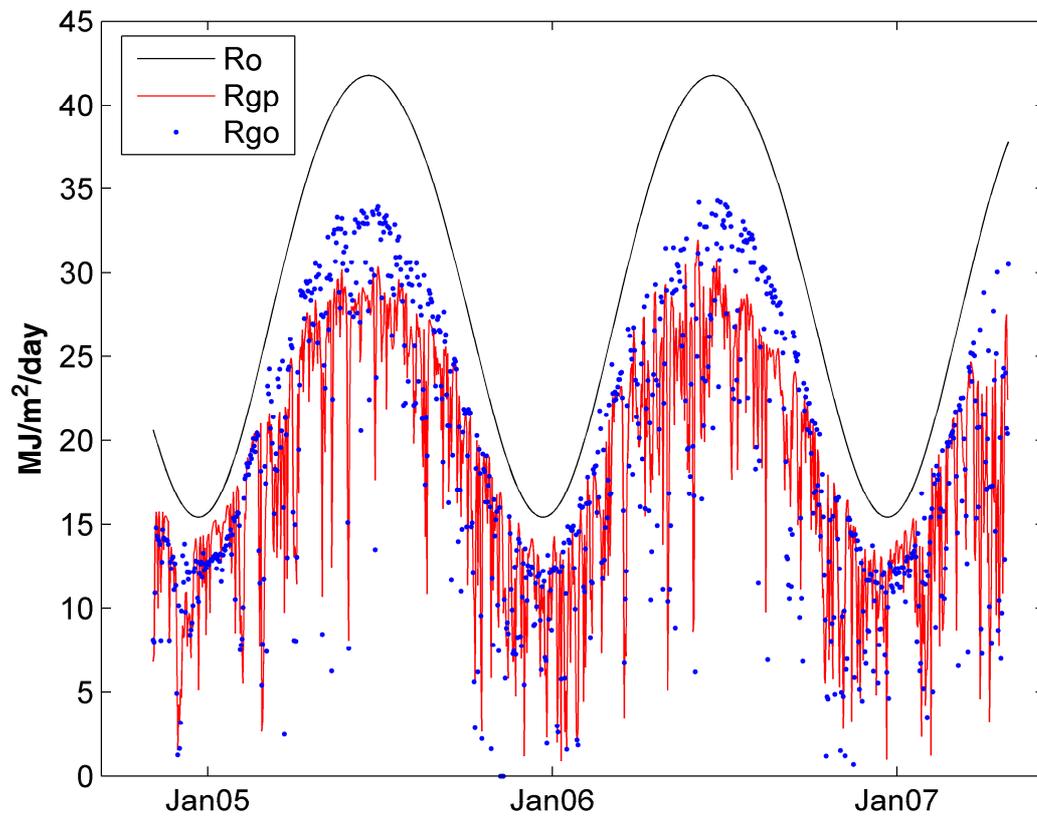


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4 Figure 5. Observed (R_{go}), and predicted (R_{gp}) global daily radiation (MJ/m²/day) at station
5 702 and 802 for the evaluation period (4/11/2004-29/04/2007)

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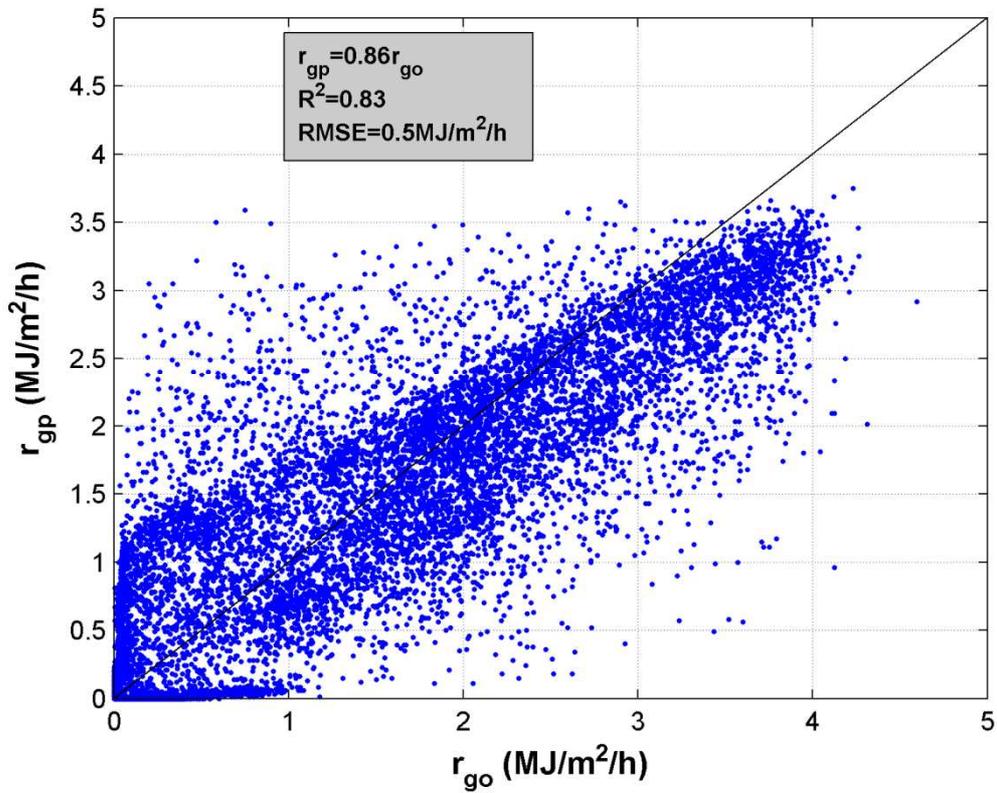
3 Figure 6. Extraterrestrial solar radiation (R_o), observed (R_{go}), and predicted (R_{gp}) global
4 radiation ($\text{MJ}/\text{m}^2/\text{day}$) at station 802 for the evaluation period (4/11/2004-29/04/2007)

5

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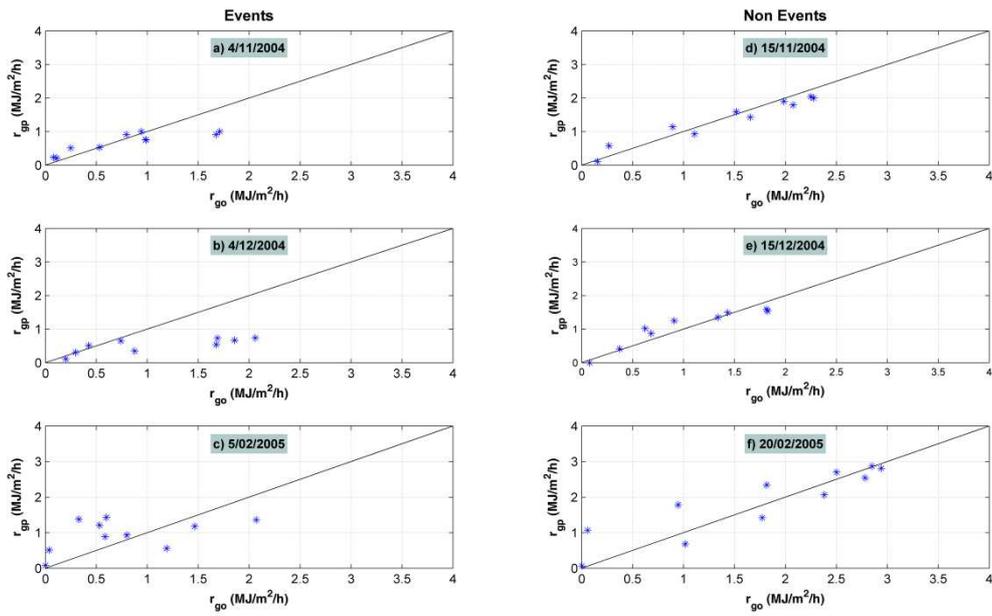
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4 Figure 7. Observed (r_{go}), and predicted (r_{gp}) global hourly radiation (MJ/m²/h) at station 802
5 for the evaluation period (4/11/2004-29/04/2007)

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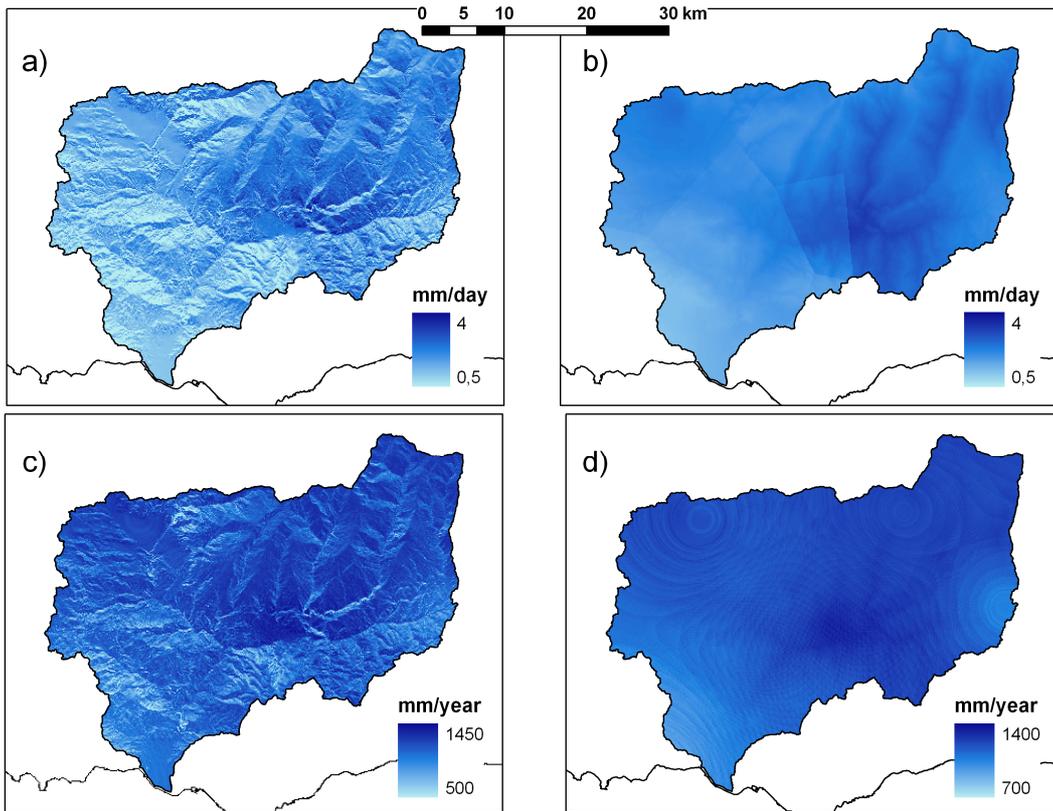


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2 Figure 8. Scatter plots of observed (r_{go}), and predicted (r_{gp}) global hourly radiation ($\text{MJ/m}^2/\text{h}$)
 3 at station 802 for certain dates

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3 Figure 9. Daily ET₀ with global radiation (20/11/2004) a) topographically corrected vs. b)

4 IDW interpolated, and annual ET₀ with global radiation (1/09/2004-31/08/2005) (c)

5 topographically corrected vs. d) IDW interpolated

6